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Quantum electrodynamics near a dispersive and absorbing dielectric

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We build up a consistent theory of quantum electrodynamics in the presence of macroscopic polarizable media. We use the Huttner-Barnett model of a dispersive and absorbing dielectric medium and formulate the theory in terms of interacting quantum fields. We integrate out the damped polaritons by using diagrammatic techniques and find an exact expression for the displacement-field (photon) propagator in the presence of a dispersive and absorbing dielectric half-space. This offers a route to traceable perturbative calculations of the same kind as in free-space quantum electrodynamics. As a worked-through example, we consider the interaction of a neutral atom with a dispersive and absorbing dielectric half-space. For that, we use the multipolar coupling \( \mathbf{\mu} \cdot \mathbf{D} \) of the atomic dipole moment to the electromagnetic displacement field. We apply this formalism to calculate the one-loop correction to the atomic electron propagator and to find the energy-level shift and changes in the spontaneous decay rates for a neutral atom close to an absorptive dielectric mirror.

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1. INTRODUCTION

Quantum electrodynamics is a well-functioning theory, which accurately predicts a wide range of phenomena not just in high-energy physics, but also in atomic physics. The best-known quantum electrodynamycal effect in atomic physics is certainly the Lamb shift, which has, by now, been calculated to very high accuracy [1]. If the atom is located not in free space but, instead, near a reflecting surface, which could be dielectric or conducting, then the reflection of photons from that surface leads to the Lamb shift acquiring an additional and, therefore, workable even for complex geometries [2].

In order to study the Casimir-Polder effect and related quantum electrodynamic effects due to the presence of macroscopic material boundaries, one needs a theory of the quantized electromagnetic field in the presence of such boundaries. The method of field quantization largely depends on how sophisticated a model of the material’s optical response one assumes. In the simplest case, one might assume perfect reflectivity of the surface. The quantization of the electromagnetic field can then be achieved quite easily by a normal-mode expansion of the field where the electromagnetic field is expanded in terms of a complete set of solutions of the homogeneous Helmholtz equation. The presence of the boundaries is taken into account by imposing appropriate boundary conditions on the electromagnetic field. Quantization is then accomplished by promoting the expansion coefficients of each mode to canonical commutation relations do not decay in time but rather take the expected form [5].

Canonical quantization of the electromagnetic field in terms of normal modes runs into difficulties when one wants to include, in the formalism, realistic properties of dielectrics. The response of the material’s surface to the electromagnetic radiation in reality depends on the frequency of the impinging radiation. Furthermore, causality requirements demand that any dispersion be always accompanied by absorption. However, a naive incorporation of absorption into canonical field quantization leads to field commutators decaying in time, i.e., an inconsistent theory. Therefore, in any model of interaction between real dielectrics and the electromagnetic field, the field has to be coupled to a reservoir in order to simulate the absorptive degrees of freedom [6,7]. This can be done in a number of ways. One is to model the absorptive degrees of freedom by adding to the operator-valued Maxwell equations Langevin-type fluctuating noise currents that ensure that the canonical commutation relations do not decay in time but rather take the expected form [8]. In this approach, the field equations are solved by using the Green’s function of the wave equation, and the noise-current operators and their properties play a major role in describing the dynamics of the coupled field-dielectric system. A few papers have provided an \textit{a posteriori} microscopic justification of such a procedure by deriving the commutative properties of the noise-current operators that were otherwise introduced \textit{ad hoc} [9–11].

A more direct approach to modeling the interaction between the electromagnetic field and an absorptive dielectric is to explicitly include, from the outset in the Lagrangian (or Hamiltonian), the matter degrees of freedom that are responsible for absorption. The dielectric is then envisaged as consisting of a continuum of harmonic oscillators coupled to a reservoir which consists of yet another set of harmonic oscillators. This quantum model of a classical dielectric was originally introduced by Hopfield [12]. The first Fano-type diagonalization [13] of the resulting Hamiltonian was achieved for fields in three dimensions in Ref. [14] for a bulk
dielectric, and the general treatment of inhomogeneous dielectrics followed in Ref. [10]. This model has also been extended to include spatial dispersion [15] and magnetodielectrics [16]. Practical applications of the Huttner-Barnett model, e.g., the calculation of spontaneous decay rates [17], work well for bulk dielectrics where simple forms of the relevant operators can be found, although an additional difficulty is that, in a bulk medium, local field corrections play an important role and need to be included. On the other hand, complications that arise due to inhomogeneities of the dielectric have previously led to unwieldy and impractical results; the conceptually very interesting paper by Yeung and Gustafson [18] uses the Wiener-Hopf method to calculate the photon propagator of the vector field \( A \) in the presence of an absorbing dielectric half-space, but the result is so complicated that it has to be Fourier transformed and evaluated numerically, whence all subsequent calculations are also necessarily only numerical.

In this paper, we demonstrate that, by starting from a Power-Zienau-Wooley type of Hamiltonian rather than adopting minimal coupling, one can carry out explicit and easy-to-follow perturbative calculations in quantum electrodynamics in the presence of inhomogeneous Huttner-Barnett dielectrics. We apply the formalism we develop to the problem of calculating the energy-level shifts and change in spontaneous-decay rates for a neutral atom placed in the vicinity of a dielectric half-space. We successfully re-derive the well-known results of phenomenological methods and broaden them by providing the asymptotic expansions that quantify the influence of absorption on the standard Casimir-Polder force calculated in Ref. [19]. We use only standard methods of quantum field theory in a similar way as this is done in condensed-matter theories. This requires the calculation of quantum propagators, most notably that of the electromagnetic field. We show that this task is nontrivial but manageable. Inspired by the results of Ref. [20], we find an exact solution of the Dyson equation satisfied by the photon propagator. In Appendix C, we make contact with the phenomenological noise-current approach and calculate the photon propagator using the electro-magnetic-field operators constructed on the basis of the noise-current operators [8].

II. CONSTRUCTION OF THE MODEL AND HAMILTONIANS

We are aiming to study the electromagnetic interaction between a quantum system, e.g., an atom, and a macroscopic absorbing dielectric body. To this end, we use the model of absorbing dielectrics developed in Ref. [6] but generalized to inhomogeneous dielectrics. The dielectric is modeled by a continuum of quantized harmonic oscillators—the polarization field. This, in turn, is coupled to another set of quantized harmonic oscillators—the reservoir, the presence of which leads to damping in the polarization field so as to allow the absorption of radiation. These coupled quantum fields interact with the electromagnetic field via the coupling of the polarization field to the electric field. It turns out that the subsystem consisting of the reservoir, the polarization, and the electromagnetic field is exactly soluble, at least, for simple geometries of the dielectric. Therefore, the interaction of the atom with the dielectric can be reduced to the interaction of the atomic dipole with the dressed electromagnetic field, that is, the electromagnetic field corrected for the presence of an absorptive body. This approach builds on the theory developed in Ref. [21] where the interaction between an atom and a pointlike absorptive dielectric (i.e., damped harmonic oscillator) was addressed. The crucial difference is that, for a description of the interaction with a pointlike absorber, the dressed electromagnetic field is required only perturbatively, but in the case of an extended absorbing body, one needs to find the dressed electromagnetic field exactly if one wants to accurately capture the interaction with an atom or other quantum system.

Our starting point is the Lagrangian density describing the complete dynamics of the electromagnetic field and the dielectric,

\[
\mathcal{L}_0 = \mathcal{L}_{\text{EM}} + \mathcal{L}_P + \mathcal{L}_R + \mathcal{L}_{\text{P-EM}}. \tag{1}
\]

The various constituent parts are as follows:

(i) The Lagrangian density \( \mathcal{L}_{\text{EM}} \) of the free electromagnetic field,

\[
\mathcal{L}_{\text{EM}} = \frac{\epsilon_0}{2} \mathbf{E}(\mathbf{r})^2 - \frac{1}{2 \mu_0} \mathbf{B}^2(\mathbf{r}). \tag{2}
\]

where \( \mathbf{E}(\mathbf{r}) \) is the electric field and \( \mathbf{B}(\mathbf{r}) \) is the magnetic induction [22].

(ii) The Lagrangian density \( \mathcal{L}_P \) of the polarization field,

\[
\mathcal{L}_P = \frac{1}{2} \mathcal{M} \mathbf{X}(\mathbf{r})^2 - \frac{1}{2} \mathcal{M}_0 \omega_0^2 \mathbf{X}(\mathbf{r})^2. \tag{3}
\]

The field \( \mathbf{X} \) is the dipole moment density of the continuum of harmonic oscillators describing the dielectric. The strength of the restoring force, acting on the polarization oscillators, is determined by the combination \( \mathcal{M} \omega_0^2 \). Hence, for a fixed absorption frequency \( \omega_0 \) of the dielectric, the mass \( \mathcal{M} \) is the parameter that determines the susceptibility of the polarization oscillator to an external agent. It has dimensions of \( (\text{mass}) \times (\text{length})^{-1} \times (\text{dipole moment density})^{-1} \). In fact, the quantity \( (\mathcal{M} \omega_0^2) \) will turn out to be the polarizability of the dielectric at zero frequency [12]. The absence of derivatives with respect to \( \mathbf{r} \) in Eq. (3) implies that the polarization oscillators at different points in space are mutually independent, resulting in a model with no spatial dispersion.

(iii) The Lagrangian density \( \mathcal{L}_R \) of the reservoir, including its coupling to the polarization field,

\[
\mathcal{L}_R = \int_0^{\infty} d\nu \left\{ \frac{1}{2} \rho_\nu \mathbf{Y}_\nu(\mathbf{r})^2 - \frac{1}{2} \rho_\nu v^2 \left[ \mathbf{Y}_\nu(\mathbf{r}) - \mathbf{X}(\mathbf{r}) \right]^2 \right\}. \tag{4}
\]

The set of fields \( \mathbf{Y}_\nu \) represents the dipole moment density of the bath oscillators at all bath frequencies \( \nu \), and the parameter \( \rho_\nu \) has dimensions of \( (\text{mass}) \times (\text{length})^{-1} \times (\text{dipole moment density})^{-2} \times (\text{frequency})^{-1} \). The coupling of the bath to the polarization field leads to the appearance of a term proportional to \( \mathbf{X}(\mathbf{r},t) \) in the equations of motion for the polarization field; hence, it is responsible for damping [23,24] (cf. also Appendix B). The masses of the bath oscillators \( \rho_\nu \) vary continuously with index \( \nu \) and describe the strength of the coupling between a single polarization oscillator and the continuum of reservoir oscillators for different frequencies \( \nu \). The precise profile of \( \rho_\nu \) is chosen so that the desired absorption spectrum is obtained [25].
(iv) The Lagrangian density \( \mathcal{L}_{P-EM} \) describing the interaction of the polarization field with the electromagnetic field,

\[
\mathcal{L}_{P-EM} = g(\mathbf{r}) \mathbf{X}(\mathbf{r}) \cdot \mathbf{E}(\mathbf{r}).
\]

The dimensionless coupling function \( g(\mathbf{r}) \) specifies where the interaction is taking place, i.e.,

\[
g(\mathbf{r}) = \begin{cases} 
1, & \text{inside the dielectric}, \\
0, & \text{outside the dielectric}.
\end{cases}
\]

Thus, \( g(\mathbf{r}) \) describes the geometric shape of the dielectric body and limits the interaction to its interior. Therefore, it is inconsequential whether the polarization field \( \mathbf{X}(\mathbf{r}) \) is defined in the whole space or only in the interior, but the latter would cause unnecessary technical complications later on.

It is straightforward to identify the canonical momenta \( \mathbf{Z}_\nu(\mathbf{r}) \equiv \rho_\nu \mathbf{Y}_\nu(\mathbf{r}), \mathbf{P}(\mathbf{r}) \equiv \mathcal{M} \mathbf{X}(\mathbf{r}), \) and \(-\mathbf{D}(\mathbf{r}) \equiv -\varepsilon_0 \mathbf{E}(\mathbf{r}) - g(\mathbf{r}) \mathbf{X}(\mathbf{r}) \) to obtain the corresponding Hamiltonian densities,

\[
\mathcal{H}_{EM} = \frac{1}{2\varepsilon_0} \mathbf{D}^2(\mathbf{r}) + \frac{1}{2\mu_0} \mathbf{B}^2(\mathbf{r}),
\]

\[
\mathcal{H}_p = \frac{\mathbf{P}(\mathbf{r})}{2\mathcal{M}} + \frac{1}{2} \mathcal{M} \omega_\nu^2 \mathbf{X}^2(\mathbf{r}),
\]

\[
\mathcal{H}_R = \frac{1}{2\rho_\nu} \int_{-\infty}^{\infty} d\nu \nu^2 (\mathbf{Z}_\nu(\mathbf{r}) - \varepsilon_0 \mathcal{M} \mathbf{Y}_\nu(\mathbf{r})),
\]

\[
\mathcal{H}_{P-R} = -\frac{g(\mathbf{r})}{\varepsilon_0} \mathbf{D}(\mathbf{r}) \cdot \mathbf{X}(\mathbf{r}).
\]

Convenience, we have separated out the polarization-field reservoir coupling \( \mathcal{H}_{P-R} \) and the part of the Hamiltonian that just shifts the eigenfrequency of the polarization field,

\[
\mathcal{H}_S = \frac{1}{2} \int_{-\infty}^{\infty} d\nu \nu^2 \mathbf{X}^2(\mathbf{r}) + \frac{1}{2\varepsilon_0} \mathbf{X}^2(\mathbf{r}).
\]

The first term of (12) arises due the coupling between the polarization field and the reservoir, whereas, the second term is caused by the coupling between the electromagnetic and the polarization fields. Equations (7)–(12), accompanied by the set of the equal-time commutation relations,

\[
[D_i(\mathbf{r}), B_j(\mathbf{r}')] = i\hbar \epsilon^{ilm} v_m(3)(\mathbf{r} - \mathbf{r}'),
\]

\[
[X_i(\mathbf{r}), P_j(\mathbf{r}')] = i\hbar \delta_{ij} \delta(3)(\mathbf{r} - \mathbf{r}'),
\]

\[
[Y_{i,v}(\mathbf{r}), Z_{j,v}(\mathbf{r}')] = i\hbar \delta_{ij} \delta(3)(\mathbf{r} - \mathbf{r}')(v - \nu'),
\]

allow one to derive the equations of motion for the inhomogeneous damped-polariton model, cf. Ref. [9]. The dielectric displacement \( \mathbf{D}(\mathbf{r}) \equiv \varepsilon_0 \mathbf{E}(\mathbf{r}) + g(\mathbf{r}) \mathbf{X}(\mathbf{r}) \) is the negative of the momentum conjugate to electromagnetic vector potential \( \mathbf{A}(\mathbf{r}) \) as it should be [26]. This is assured by the correct choice of coupling (5). As already mentioned, the Hamiltonian density \( \mathcal{H}_S \), Eq. (12), shifts the eigenfrequency \( \omega_\tau \) of the polarization field, i.e.,

\[
\omega^2_\tau \longrightarrow \tilde{\omega}^2_\tau = \omega_\tau^2 + \frac{1}{\mathcal{M}} \int_{\text{volume}} d\nu \rho_\nu \nu^2 + \frac{g^2(\mathbf{r})}{\varepsilon_0 \mathcal{M}}.
\]

The second term contains the parameter \( \rho_\nu \), that pertains to the shape of the absorption spectrum. For our choice of \( \nu \) dependence (see Appendix B), it turns out to be infinite. However, this is not problematic as the equations of motion for the fields, and hence, all observable quantities, most notably the dielectric function, stay finite and physically meaningful. Furthermore, the last term of Eq. (16), in principle, introduces a position dependence of the frequency \( \omega_\tau \) through the coupling function \( g(\mathbf{r}) \). While not yet apparent at this stage, this position dependence will turn out to be irrelevant. Hence, for now, we set \( g(\mathbf{r}) = 1 \) in the expression for the frequency shift but will explain later on why we are allowed to do so. With that, we can incorporate Eq. (12) into the Hamiltonian density of the polarization field and can write

\[
\mathcal{H}_p = \frac{\mathbf{P}^2(\mathbf{r})}{2\mathcal{M}} + \frac{1}{2} \mathcal{M} \omega_\tau^2 \mathbf{X}^2(\mathbf{r}),
\]

with

\[
\tilde{\omega}^2_\tau = \omega^2 + \omega_\tau^2 + \frac{1}{\mathcal{M}} \int_{\text{volume}} d\nu \rho_\nu \nu^2,
\]

where by hindsight, we have introduced the symbol \( \omega_\nu^2 = (\varepsilon_0 \mathcal{M})^{-1} \) in analogy to the plasma frequency in metals [27].

Our aim is to investigate the influence of an absorbing dielectric on the properties of an atom, such as the shifts in its energy levels and spontaneous decay rates. The complete Hamiltonian, including the atom, reads

\[
H = \int d^3 r [\mathcal{H}_A + \mathcal{H}_{EM} + \mathcal{H}_p + \mathcal{H}_R + \mathcal{H}_{P-R} + \mathcal{H}_{P-EM} + \mathcal{H}_{A-EM}],
\]

where \( \mathcal{H}_A \) is the Hamiltonian density of the atom and \( \mathcal{H}_{A-EM} \) describes its coupling to the electromagnetic field. We consider a one-electron atom and treat the atomic electron nonrelativistically by representing it as a quantum of the Schrödinger field satisfying fermionic anticommutation relations. The Hamiltonian density \( \mathcal{H}_A \) of the noninteracting atomic electron (i.e., in the absence of interactions with the quantized electromagnetic field) can be written as

\[
\mathcal{H}_A = \Psi^\dagger(\mathbf{r}) \left[ -\frac{\hbar^2}{2m} \nabla^2 + V(\mathbf{r} - \mathbf{R}) \right] \Psi(\mathbf{r}),
\]

where \( \Psi(\mathbf{r}) \) is the Schrödinger field operator satisfying the anticommutation relation,

\[
\{\Psi(\mathbf{r}), \Psi^\dagger(\mathbf{r}')\} = \delta(3)(\mathbf{r} - \mathbf{r}'),
\]

and \( V(\mathbf{r} - \mathbf{R}) \) is the potential due to the immobile nucleus which we choose to be located well outside the dielectric (i.e., at least a few Bohr radii away) at a position \( \mathbf{R} \) so that there is no wave-function overlap between the atom and the solid. The atom is coupled to the electric field via its electric dipole moment. The Hamiltonian, describing this coupling in the dipole approximation, may be written as

\[
\mathcal{H}_{A-EM} = -\mathbf{\mu} \cdot \mathbf{E}(\mathbf{R}),
\]

i.e., we evaluate the electric field at the position of the nucleus. Here, \( \mathbf{\mu} \) is the electric dipole moment operator, which depends on the second-quantized fields \( \Psi \) and \( \Psi^\dagger \). It is convenient to expand the field operator \( \Psi(\mathbf{r}) \) in terms of a complete set of
atomic wave functions satisfying

$$\left[ -\frac{\hbar^2}{2m} \nabla^2 + V(|\mathbf{r} - \mathbf{R}|) \right] \phi_n(r) = E_n \phi_n(r).$$  \hspace{1cm} (23)

If we write

$$\Psi(r) = \sum_m c_m \phi_m(r), \quad \Psi^\dagger(r) = \sum_m c_m^\dagger \phi_m^*(r),$$  \hspace{1cm} (24)

then it follows from Eq. (21) that the operators $c_m$ and $c_m^\dagger$ satisfy the equal-time anticommutation relation,

$$\{c_n, c_m^\dagger\} = \delta_{mn}.$$  \hspace{1cm} (25)

We use Eqs. (24) and (25) to rewrite the Hamiltonians $H_A$ and $H_{EM}$ in a more useful form,

$$H_A = \sum_n E_n c_n^\dagger c_n,$$  \hspace{1cm} (26)

$$H_{A-EM} = -e \sum_{ij} c_i^\dagger c_j \langle i| \rho | j \rangle \cdot \mathbf{E}(\mathbf{R}),$$  \hspace{1cm} (27)

where $-e \langle i| \rho | j \rangle$ are the dipole matrix elements.

We will follow the field-theoretical approach of Ref. [21] to calculate the energy-level shifts and modified spontaneous decay rates. In order to do so, we need to locate the poles of the atomic propagator, which, once interactions have been switched on, are accessible only perturbatively. For these perturbative calculations, we need to work in the interaction picture where the general expression for the perturbative expansion of a Green’s function of the field $\Psi$ under the influence of the interaction $H_I$ is [28]

$$G_{\text{conn}}(\mathbf{r}, \mathbf{r}', t, t') = \sum_{n=0}^{\infty} \left( -\frac{i}{\hbar} \right)^n \int dt_1 \cdots \int dt_n$$

$$\times \langle \Omega | \Psi(\mathbf{r}, t) \Psi^\dagger(\mathbf{r}', t') H_I(t_1) \cdots H_I(t_n) | \Omega \rangle_{\text{conn}},$$  \hspace{1cm} (28)

$\Psi$ is now the field operator in the Heisenberg picture, and $| \Omega \rangle$ is the exact ground state of the noninteracting system. The subscript “conn” indicates that only connected diagrams contribute as disconnected diagrams drop out in the normalization of $| \Omega \rangle$.

Wick’s theorem states that the terms appearing in the expansion (28), when written out explicitly for a specific interaction Hamiltonian, turn out to be given entirely in terms of the propagators of the noninteracting fields. We will proceed by determining the noninteracting propagators of the atom, the polarization field, the bath, and the electromagnetic field. Then, the interaction of the polarization field with the reservoir will be treated exactly to all orders. Once this is accomplished, the correction to the electromagnetic-field propagator caused by the presence of the absorbptive dielectric can be calculated, which will also be done exactly to all orders. This is going to give the dressed photon propagator that enters the final perturbative expansion of the atomic propagator whose poles give the energy-level shifts and changes in the transition rates.

### III. UNPERTURBED FEYNMAN PROPAGATORS

#### A. Atomic-electron propagator

The unperturbed atomic-electron propagator, corresponding to the Hamiltonian (20) or equivalently (26), is defined as the time-ordered expectation value,

$$G^{(0)}(\mathbf{r}, \mathbf{r}', t, t') = -\frac{i}{\hbar} \langle \Omega | \Psi(\mathbf{r}, t) \Psi^\dagger(\mathbf{r}', t') | \Omega \rangle,$$  \hspace{1cm} (29)

where $\Psi$ is the Schrödinger field operator in the Heisenberg picture and $| \Omega \rangle$ is the exact ground state of the noninteracting system. We substitute the field operators written in terms of the atomic eigenfunctions, Eq. (24), while remembering that we are in the Heisenberg picture where the operators $c_l$ and $c_l^\dagger$ are time dependent and find

$$G^{(0)}(\mathbf{r}, \mathbf{r}', t, t') = \sum_{lj} \phi_l(\mathbf{r}) \phi^*_l(\mathbf{r}') G^{(0)}_{lm}(t, t'),$$  \hspace{1cm} (30)

with

$$G^{(0)}_{lm}(t, t') = -\frac{i}{\hbar} \langle \Omega | \Psi(\mathbf{r}, t) c_m^\dagger(t') | \Omega \rangle.$$  \hspace{1cm} (31)

The time dependence of the fermionic annihilation and creation operators is governed by the Hamiltonian (26),

$$c_m(t) = c_m(0) e^{-i E_m t / \hbar}, \quad c_m^\dagger(t) = c_m^\dagger(0) e^{i E_m t / \hbar}.$$  \hspace{1cm} (32)

With that, we can determine $G^{(0)}_{lm}(t, t')$ and obtain

$$G^{(0)}_{lm}(t - t') = -\frac{i}{\hbar} \delta(t - t') e^{-i E_m(t - t') / \hbar} \delta_{lm},$$  \hspace{1cm} (33)

where we have used the definition of the time-ordering operator and the fact that the vacuum state $| \Omega \rangle$ is annihilated by $c_m(0)$. Since $G^{(0)}_{lm}(t, t')$ is, in fact, dependent only on the time difference $t - t'$, we can work with its Fourier transform with respect to $t - t' \equiv \tau$,

$$G^{(0)}_{lm}(E) = \int_{-\infty}^{\infty} d\tau \ e^{i E \tau / \hbar} G^{(0)}_{lm}(\tau) = \frac{1}{E - E_l} \delta_{lm}.$$  \hspace{1cm} (34)

With this convention of Fourier transformation, the $i \hbar$ prescription ensures the correct causal behavior of the propagator and guarantees the convergence of the integrals.

#### B. Photon propagator

To calculate the zeroth-order propagator of the displacement field $\mathbf{D}(\mathbf{r}, t)$ whose dynamics is governed by the Hamiltonian (7), which we emphasize does not include the coupling term (11), we note that the Heisenberg equations of motion imply

$$\frac{\partial D_l(\mathbf{r}, t)}{\partial t} = \frac{1}{\mu_0} \epsilon^{ijk} \nabla_j B_k(\mathbf{r}, t),$$  \hspace{1cm} (35)

$$\frac{\partial B_l(\mathbf{r}, t)}{\partial t} = -\frac{1}{\epsilon_0} \epsilon^{ijk} \nabla_j D_k(\mathbf{r}, t),$$  \hspace{1cm} (36)

where $\epsilon^{ijk}$ is the Levi-Civita symbol and the sum over doubly occurring Cartesian indices is implied. Thus, the displacement
field $\mathbf{D}(\mathbf{r},t)$ satisfies the homogeneous wave equation,

$$(\nabla_j \nabla_j - \delta_{ij} \nabla^2) D_j(\mathbf{r},t) + \mu_0 \varepsilon_0 \frac{\partial^2}{\partial t^2} D_i(\mathbf{r},t) = 0. \tag{37}$$

The formal definition of the photon propagator reads

$$D_{ij}(\mathbf{r},\mathbf{r}',t,t') = -\frac{i}{\hbar} \left[ T[D_i(\mathbf{r},t)D_j(\mathbf{r}',t')] \right] \tag{38}$$

where $D_i(\mathbf{r},t)$ is the displacement-field operator in the Heisenberg picture and $|0\rangle$ is the exact ground state of the noninteracting electromagnetic field. We proceed by applying the differential wave operator that appears in Eq. (37) to this definition of the propagator, but we need to take care when applying the time derivative to a time-ordered product and observe that

$$\frac{\partial}{\partial t} T[A(t)B(t')] = \delta(t-t')[A(t), B(t')] + T \left[ \frac{\partial A(t)}{\partial t} B(t') \right].$$

Thus, we find that the displacement-field propagator $D^{(0)}_{ij}(\mathbf{r},\mathbf{r}',t,t')$ satisfies the following differential equation:

$$\left(\nabla_i \nabla_j - \delta_{ij} \nabla^2 + \mu_0 \varepsilon_0 \frac{\partial^2}{\partial t^2} \right) D^{(0)}_{ij}(\mathbf{r} - \mathbf{r}', t - t') = \frac{\varepsilon_0}{(2\pi)^2} \delta(t-t') \int d^3 \mathbf{q} (q_i q_k - \delta_{ik} q^2) e^{i\mathbf{q} \cdot (\mathbf{r} - \mathbf{r}')} \tag{39}$$

and the fact that spatial derivatives commute with the time-ordering operator. From Eq. (39), it is clear that the free-space photon propagator is translation invariant in space and time, i.e., it depends only on the differences $\mathbf{r} - \mathbf{r}'$ and $\tau = t - t'$. Therefore, one can find the solution of the differential equation through Fourier transformation. First, we note that Maxwell’s equation (35) implies that the displacement field is transverse so that its propagator satisfies

$$\nabla_i D^{(0)}_{ij}(\mathbf{r} - \mathbf{r}', t - t') = 0. \tag{40}$$

Introducing the Fourier transform of the propagator,

$$D^{(0)}_{ik}(\mathbf{q},\omega) = \int d^3 \mathbf{r} e^{-i\mathbf{q} \cdot \mathbf{r}} D^{(0)}_{ik}(\mathbf{r} - \mathbf{r}', t - t') 
\times \int_{-\infty}^{\infty} d(t-t')e^{i\omega(t-t')} D^{(0)}_{ik}(\mathbf{r} - \mathbf{r}', t - t'), \tag{41}$$

we readily obtain its spectral representation,

$$D^{(0)}_{ik}(\mathbf{q},\omega) = \varepsilon_0 \delta_{ik} \left[ \mathbf{q}^2 - q_i q_k \right] \frac{\delta(\omega - \omega_0)}{\omega^2 - \mathbf{q}^2 + \frac{i\hbar}{2}}. \tag{42}$$

We have displaced the poles in the denominator by $i\eta$ so that $D^{(0)}_{ik}(\mathbf{r} - \mathbf{r}', t - t')$ has the appropriate causality properties of a Feynman propagator.

C. Polarization-field propagator

The Hamiltonian density (17) describes a collection of mutually independent harmonic oscillators. The fact that the harmonic oscillator at $\mathbf{r}$ is unaffected by the oscillator at $\mathbf{r} + d\mathbf{r}$ allows us to introduce creation and annihilation operators $b^\dagger(\mathbf{r})$ and $b(\mathbf{r})$ for each harmonic oscillator,

$$X(\mathbf{r}) = \sqrt{\frac{\hbar}{2\omega_0}}[b^\dagger(\mathbf{r}) + b(\mathbf{r})], \tag{43}$$

$$P(\mathbf{r}) = i\sqrt{\frac{\hbar \omega_0}{2}}[b^\dagger(\mathbf{r}) - b(\mathbf{r})].$$

The operators $b^\dagger(\mathbf{r})$ and $b(\mathbf{r})$ satisfy the equal-time commutation relations,

$$[b_i(\mathbf{r}), b^\dagger_j(\mathbf{r})] = \delta_{ij} \delta^{(3)}(\mathbf{r} - \mathbf{r}'), \tag{44}$$

which follow directly from their definition and Eq. (15). The operator $b_i(\mathbf{r})$ annihilates the ground state of the oscillation in the $i$th direction at $\mathbf{r}$. Using this property together with the commutation relation (44), we can directly evaluate the polarization-field propagator defined as

$$K^{(0)}_{ij}(\mathbf{r},\mathbf{r}',t,t') = -\frac{i}{\hbar} \langle \Omega | T[X_i(\mathbf{r},t)X_j(\mathbf{r}',t')] | \Omega \rangle. \tag{45}$$

Here, $X_i(\mathbf{r},t)$ is the polarization-field operator in the Heisenberg picture, and $|\Omega\rangle$ is the exact ground state of the noninteracting polarizing field. When written in terms of the creation and annihilation operators, the Hamiltonian density (17) is, of course, diagonal in $b^\dagger(\mathbf{r})$ and $b(\mathbf{r})$ so that the time dependence of the creation and annihilation operators is harmonic,

$$b_i(\mathbf{r},t) = b_i(\mathbf{r},0)e^{-i\omega t}, \quad b^\dagger_i(\mathbf{r},t) = b^\dagger_i(\mathbf{r},0)e^{i\omega t}. \tag{46}$$

We substitute the polarization-field operators (43) expressed in terms of the ladder operators into Eq. (45) and observe that, due to the orthogonality of states, only terms proportional to $b_i b^\dagger_i$ contribute. Taking care of the appropriate time ordering of operators and using the commutator (44) to move any annihilation operators to the right of creation operators so that they act on the vacuum state $|\Omega\rangle$, we readily obtain

$$K^{(0)}_{ij}(\mathbf{r} - \mathbf{r}',t-t') = -\frac{i}{2\omega_0} \delta^{(3)}(\mathbf{r} - \mathbf{r}')$$

$$\times [\theta(t-t')e^{-i\omega(t-t')} + \theta(t'-t)e^{i\omega(t-t')}], \tag{47}$$

with the frequency $\omega_0$ as defined in Eq. (18). We will need the Fourier transform of the polarization propagator with respect to the time difference $t - t'$,

$$K^{(0)}_{ij}(\mathbf{r} - \mathbf{r}';\omega) = \int_{-\infty}^{\infty} d(t-t')e^{i\omega(t-t')} K^{(0)}_{ij}(\mathbf{r} - \mathbf{r}',t-t'), \tag{48}$$

which is easily obtained from Eq. (47) and reads

$$K^{(0)}_{ij}(\mathbf{r} - \mathbf{r}';\omega) = \frac{1}{\omega_0^2 - \omega^2 + \frac{i\hbar}{2}} \delta^{(3)}(\mathbf{r} - \mathbf{r}'). \tag{49}$$

Since the polarization-field operators are mutually independent, there is no need for any special consideration of the boundaries of the dielectric medium at this stage. The boundaries are taken into account through the coupling function.
g(r), Eq. (6), and the field equation for the electromagnetic field, once coupled to the polarization field, will include the physical processes of reflection and refraction as it should. It is worth pointing out that an artificial restriction of the free polarization field to just the interior of the dielectric would lead to a much more complicated free propagator thereby causing unnecessary technical complications while not describing any different physics.

D. Reservoir propagator

The dynamics of the noninteracting reservoir field is governed by the Hamiltonian (9), which describes a set of independent harmonic oscillators. The propagator for the free reservoir field can be obtained by repeating the same steps as for the derivation of the propagator for the free polarization field in Sec. III C. Therefore, we do not repeat the details of the derivation but simply point out the similarity of the structure of the result to Eqs. (47) and (49). In the time domain, the reservoir propagator reads

\[ H_{ij}^{(0)}(r, r', t, t') = -\frac{i}{\hbar} e^{-i\delta_{ij}(r - r')/\hbar} \delta(v - v') \times [\theta(t - t')e^{-iv(t - t')} + \theta(t' - t)e^{iv(t - t')}. \] (50)

Its Fourier transform, with respect to \( t - t' \), is given by

\[ H_{ij}^{(0)}(r, r', v, v', \omega) = \int_{-\infty}^{\infty} d(t - t')e^{i\omega(t - t')} H_{ij}^{(0)}(r, r', t - t', v, v'), \]

\[ = \frac{1}{\hbar} \frac{\delta_{ij}}{\omega^2 - \nu^2 - i\eta} \delta^{(3)}(r - r')\delta(v - v'). \] (51)

IV. DRESSED PROPAGATORS

Having gathered all the unperturbed propagators, we can proceed to work out the propagators for the coupled fields. We are going to use a diagrammatic approach to illustrate the workings of perturbation theory, i.e., we represent each term of the perturbative expansion in Eq. (28) by an appropriate Feynman diagram (cf., e.g., Ref. [28]). To proceed with that, we need to lay down the Feynman rules for our approach. We have four different free propagators; accordingly, we associate four distinct lines with them:

We will need to consider three interaction Hamiltonians, in turn, \( H_{p-r} \), \( H_{p-em} \), and \( H_{a-em} \).

\[ H_{p-r} = -\int d^3r \int_0^\infty dv \rho_0v^2 X(r) \cdot Y(r), \] (52)

\[ H_{p-em} = -\frac{1}{\epsilon_0} \int d^3r g(r) D(r) \cdot X(r), \] (53)

\[ H_{a-em} = -\frac{1}{\epsilon_0} \sum_{ij} c_j^\dagger c_j \mu_{ij} \cdot D(R). \] (54)

Note that we have introduced the shorthand \( \mu_{ij} = (i|\mu|j) \) for the matrix elements of the atomic electric dipole moment operator \( \mu \). These interaction Hamiltonians yield the following Feynman rules for the vertices between the lines defined above:

To compute a diagram, one has to sum over all internal indices and to integrate over internal times, internal coordinates, and reservoir oscillator frequencies \( v \) and \( v' \). As mentioned earlier, the subscript conn in Eq. (28) means that the summation in that equation runs only over those terms that correspond to connected Feynman diagrams. Furthermore, topologically equivalent diagrams, i.e., those that can be obtained by permuting the factors \( H_i(t_i) \) in Eq. (28), are counted only once, and therefore, we have omitted the factor of \( 1/n! \) that would otherwise have arisen in a straightforward expansion of the time-ordered exponential of the interaction Hamiltonian in perturbation theory.

A. Dressing the polarization line

The polarization field interacts with the reservoir; all these interactions in their entirety dress the polarization field. We choose to represent the dressed polarization propagator by a bold dashed line:

From the interaction Hamiltonian (52) and the associated Feynman rules, one can see that the polarization line can only ever connect to exactly one reservoir line. Hence, the complete set of all possible interactions, corresponding to the expansion (28), is represented by the following sequence of Feynman
The equivalent analytical expression is the Dyson equation for the dressed polarization propagator; it reads
\[
K_{mn}(\mathbf{r}, \mathbf{r}', t, t') = K^{(0)}_{mn}(\mathbf{r}, \mathbf{r}', t, t')
+ \sum_{l,p} \int_{-\infty}^{\infty} dt_1 \int_{-\infty}^{\infty} dt_2 \int d^3 \mathbf{r}_1 \int d^3 \mathbf{r}_2 \int_0^{\infty} dv \int_0^{\infty} dv' \times K^{(0)}_{ml}(\mathbf{r}_1, t, t_1) H^{(0)}_{lp}(\mathbf{r}_1, \mathbf{r}_2, t_1, t_2, v, v') K_{pm}(\mathbf{r}_2, \mathbf{r}', t_2, t').
\]

Despite being an integral equation, the above equation is easily solved exactly. Upon substituting Eqs. (49) and (51), we can easily carry out the spatial integrations. Then, we Fourier transform with respect to \( t - t' \),
\[
K_{mn}(\mathbf{r}, \mathbf{r}'; \omega) = \int_{-\infty}^{\infty} dt (t - t') e^{i\omega(t-t')} K_{mn}(\mathbf{r}, \mathbf{r}', t - t'),
\]
and find the following expression for the dressed polarization-field propagator in the frequency domain,
\[
K_{mn}(\mathbf{r} - \mathbf{r}'; \omega) = K(\omega) \delta^{(3)}(\mathbf{r} - \mathbf{r}') \delta_{mn},
\]
with
\[
K(\omega) = \frac{1}{\mathcal{M}} \left[ \omega^2 - \omega_T^2 - \omega_p^2 - \frac{\omega^2}{\mathcal{M}} \int_{-\infty}^{\infty} dv \left( 1 - \frac{\rho \nu^2}{\omega^2 - \nu^2 + i\eta} \right) \right]^{-1}.
\]

Note that \( K(\omega) \) is an even function of \( \omega \). The plasma frequency \( \omega_p \) was defined below Eq. (18).

### B. Dressing the photon line

The coupling (53) between the dressed polarization field and the electromagnetic field has formally the same form as the coupling (52) between the bare polarization field and the reservoir. By analogy with the previous section, we write down the graphical equation for the dressed photon propagator as
\[
\begin{align*}
\begin{array}{c}
\vdots
\end{array}
\end{align*}
\]

where the bold wavy line denotes the dressed photon propagator, i.e.,
\[
D_{ik}(\mathbf{r}, \mathbf{r}', t, t') = e^{i\mathcal{H}t} \mathcal{D}_{kl}(\mathbf{r}, \mathbf{r}', t, t')
\]

The corresponding analytical expression reads
\[
D_{ik}(\mathbf{r}, \mathbf{r}', t, t') = D^{(0)}_{ik}(\mathbf{r}, \mathbf{r}', t, t')
+ \frac{1}{\epsilon_0} \sum_{j,l} \int_{-\infty}^{\infty} dt_1 \int_{-\infty}^{\infty} dt_2 \int d^3 \mathbf{r}_1 \int d^3 \mathbf{r}_2 g(\mathbf{r}_1) g(\mathbf{r}_2) \times D^{(0)}_{jl}(\mathbf{r} - \mathbf{r}_1, t - t_1) K_{jl}(\mathbf{r}_1 - \mathbf{r}_2, t_1, t_2) D_{kl}(\mathbf{r}_2, \mathbf{r}', t_2, t').
\]

Now, recall the discussion following Eq. (16) of the shifted eigenfrequency \( \tilde{\omega}_T \) of the polarization field. It enters the Dyson equation (57) through the dressed polarization-field propagator \( K_{jl}(\mathbf{r}_1 - \mathbf{r}_2, t_1, t_2) \). As we noted earlier, according to Eq. (16), the shifted eigenfrequency \( \tilde{\omega}_T \) suddenly jumps at the boundary of the region where the polarization field interacts with the electromagnetic field, i.e., where the coupling function \( g(\mathbf{r}) = 1 \). However, it is now apparent that this discontinuity is unproblematic because all spatial integrations in Eq. (57) are limited to the region of space where \( g(\mathbf{r}) = 1 \).

To simplify Eq. (57), we note that one of the spatial integrations is trivial due to the \( \delta \) function in the dressed polarization-field propagator (55). Fourier transforming with respect to the time difference \( t - t' \),
\[
D_{ik}(\mathbf{r}, \mathbf{r}', \omega) = \int_{-\infty}^{\infty} dt (t - t') e^{i\omega(t-t')} D_{ik}(\mathbf{r}, \mathbf{r}', t, t'),
\]
we find the Dyson equation for the dressed photon propagator in the frequency domain,
\[
D_{ik}(\mathbf{r}, \mathbf{r}', \omega) = D^{(0)}_{ik}(\mathbf{r} - \mathbf{r}', \omega) + \frac{K(\omega)}{\epsilon_0} \times \int d^3 \mathbf{r}_1 g(\mathbf{r}_1) D^{(0)}_{ij}(\mathbf{r} - \mathbf{r}_1; \omega) D_{jk}(\mathbf{r}_1, \mathbf{r}', \omega).
\]

Here, \( K(\omega) \) is a complex-valued function of frequency that has originated from the dressed polarization-field propagator and is given in Eq. (56); it will be shown to be related to the dielectric permittivity. Note that the dimensionless coupling function \( g(\mathbf{r}) \), describing the geometry of the dielectric medium as defined in Eq. (6), is the only way the geometry enters in the calculation by effectively defining the limits of the spatial integration in Eq. (59). \( D^{(0)}_{ik}(\mathbf{r} - \mathbf{r}', \omega) \) is the free-space photon propagator in coordinate representation, i.e., the inverse Fourier transform of Eq. (42),
\[
D^{(0)}_{ik}(\mathbf{r} - \mathbf{r}', \omega) = \frac{\epsilon_0}{(2\pi)^3} \int d^3 q e^{i\mathbf{q}\cdot\mathbf{r}-i\mathbf{q}\cdot\mathbf{r}'} \delta_{ik} \frac{q^2 - q_1 q_2}{\omega^2 - q^2 + i\eta}.
\]

The solution of the integral equation (59) is much less trivial than that of the equivalent equation for dressing the polarization line in Sec. IV A. This is because translation invariance is lost when an inhomogeneous dielectric is introduced into the system. Here, we report two ways of tackling the problem. First, we demonstrate that it is possible to solve the integral equation (59) by direct iteration. The iteration method that we employ is inspired by Ref. [20]. In order to explain it, we write Eq. (59) symbolically as
\[
D = D^0 + K D^0 \otimes D.
\]

Iteration of this equation yields the expansion,
\[
D = D^0 + K D^0 \otimes D^0 + K^2 D^0 \otimes D^0 \otimes D^0 + \cdots,
\]
which proves especially useful if the action of the operator \( \mathcal{O} = D^0 \otimes \mathcal{O} \) on the free-space propagator \( D^0 \) amounts to a simple multiplication, i.e., if
\[
\mathcal{O} D^0 = D^0 \otimes D^0 = C D^0,
\]
corresponds to treating the electromagnetic field should not shows that some components of the Fourier transform of and then Fourier transform of which we know how to sum to all orders. An alternative approach, which we sketch in Appendix C, consists of converting the integral equation (59) to a differential equation supplemented by Maxwell boundary conditions. In addition, in Appendix C, for comparison with other theories, we construct the photon propagator using yet another, completely different method based on the phenomenological noise-current approach of Ref. [8].

Let us now concentrate on the example geometry of a dielectric half-space occupying the \( z < 0 \) region of space, cf. Fig. 1, for which the coupling function \( g(\mathbf{r}) \) in Eq. (59) becomes \( g(\mathbf{r}) = \theta(-z) \). Due to the boundary, the problem has lost translation invariance in the \( z \) direction but not in directions parallel to the surface. In other words, the propagator depends only on the difference \( \mathbf{r}_i - \mathbf{r}_i' \) but separately on \( z \) and \( z' \). It is convenient to work with quantities that have been Fourier transformed with respect to \( \mathbf{r}_i - \mathbf{r}_i' \); e.g., for the dressed photon propagator, we have

\[
D_{ij}(z,z') = \int d^2(\mathbf{r}_i - \mathbf{r}_i') e^{-iq_i(\mathbf{r}_{ij} - \mathbf{r}_{ij}')} D_{ij}(\mathbf{r}_i - \mathbf{r}_i', z,z'),
\]

where, for notational convenience, we have suppressed the dependence on \( \mathbf{q}_i \) and \( \omega \). Once Fourier transformed with respect to \( \mathbf{r}_i - \mathbf{r}_i' \), the integral equation (59) becomes

\[
D_{ik}(z,z') = D_{ik}^{(0)}(z - z') + \frac{K(\omega)}{\varepsilon_0} \int_0^\infty d\tau D_{ij}^{(0)}(z - z_1)D_{jk}(z_1,z').
\]

As is easily checked, this does not lend itself to iteration as it stands. Following Ref. [20], we introduce an additional integral equation in order to enable the iteration process,

\[
D_{ik}(z,z') = D_{ik}^{(c)}(z - z') - \frac{K(\omega)}{\varepsilon_0} \int_0^\infty d\tau D_{ij}^{(c)}(z - z_1)D_{jk}(z_1,z').
\]

Here, \( D_{ik}^{(c)}(z - z') \) is the Fourier-transformed photon propagator in a bulk medium, i.e., the solution of Eq. (59) with \( g(\mathbf{r}) = 1 \). In order to justify Eq. (67), let us recall that the part of the Hamiltonian density that describes the interaction of the photon field with the polarization field has the form

\[
\mathcal{H} = \mathcal{H}_0 - \frac{\theta(-z)}{\varepsilon_0} \mathbf{X}(\mathbf{r}) \cdot \mathbf{D}(\mathbf{r}),
\]

where \( \mathcal{H}_0 \) is the Hamiltonian density of the noninteracting electromagnetic field. Using the fact that \( \theta(-z) + \theta(z) = 1 \), we can also write

\[
\mathcal{H} = \mathcal{H}_e + \frac{\theta(z)}{\varepsilon_0} \mathbf{X}(\mathbf{r}) \cdot \mathbf{D}(\mathbf{r}),
\]

where \( \mathcal{H}_e = \mathcal{H}_0 - \mathbf{X}(\mathbf{r}) \cdot \mathbf{D}(\mathbf{r})/\varepsilon_0 \) is the Hamiltonian density of the electromagnetic field interacting with an unbounded dielectric. Therefore, we have a choice: We can either correct the free-space photon propagator for the presence of the dielectric half-space or, equivalently, correct the photon propagator in a bulk dielectric for the absence of the medium in the other half-space. In other words, it is entirely up to us which Hamiltonian we take as the zeroth-order (exactly solvable) Hamiltonian when applying perturbation theory. The integral equation (67) corresponds to treating the electromagnetic field interacting with the bulk medium as the zeroth-order, solved part of the problem.

To proceed, we need to find \( D_{ik}^{(0)}(z - z') \) and \( D_{ik}^{(c)}(z - z') \) appearing in Eqs. (66) and (67), which can, in fact, be read off from the representations of these propagators as two-dimensional integrals over the momenta parallel to the surface. To find \( D_{ik}^{(0)}(z - z') \), we carry out the \( q_i \) integral in Eq. (60) using the residue theorem. The result, written in a compact form, is

\[
D_{ik}^{(0)}(\mathbf{r} - \mathbf{r}', \omega) = -\frac{i\varepsilon_0}{2(2\pi)^2} \left( \nabla \cdot \mathbf{v}_k - \delta_{ik} \nabla^2 \right) \times \int d^2q_i \epsilon^{q_i(\mathbf{r}_i - \mathbf{r}_i')} e^{i|\mathbf{q}|(z - z')} k_z,
\]

where \( k_z \) is the \( z \) component of the wave vector in vacuum and is given by \( k_z = \sqrt{\omega^2 - q_i^2} + i\eta \). The square root is taken such that the imaginary part of \( k_z \) is always positive. Equation (70) shows that some components of the Fourier transform of the free-space photon propagator are singular when crossing the \( z = z' \) plane.

For deriving the photon propagator \( D_{ik}^{(c)}(\mathbf{r} - \mathbf{r}' ; \omega) \) in a bulk medium, we set \( g(\mathbf{r}) = 1 \) in Eq. (59) and then Fourier transform this equation with respect to \( \mathbf{r} - \mathbf{r}' \),

\[
D_{ik}^{(c)}(\mathbf{q}, \omega) = \frac{\varepsilon_0}{\omega^2 - \mathbf{q}^2 + i\eta} D_{jk}(\mathbf{q}, \omega).
\]

This matrix equation becomes an algebraic one when one takes the transversality of the propagator \( q_j D_{jk}^{(c)}(\mathbf{q}, \omega) = 0 \) into account. The calculation is straightforward, and in coordinate space, we obtain

\[
D_{ik}^{(c)}(\mathbf{r} - \mathbf{r}', \omega) = \frac{\varepsilon_0\xi(\omega)}{(2\pi)^3} \int d^3q \frac{\delta_{ik} \mathbf{q}^2 - q_i q_j}{\xi(\omega)\omega^2 - \mathbf{q}^2} e^{i\mathbf{q}(\mathbf{r} - \mathbf{r}')}. \]

Note that the function \( \xi(\omega) \) that appears in Eq. (72) should not be interpreted as the dielectric function of the bulk medium.
It is an even function of the frequency $\omega$ and may be written explicitly as
\[
\xi(\omega) = \left(1 + \frac{K(\omega)}{\epsilon_0}\right)^{-1} = 1 + \frac{1}{\epsilon_0 M} \left[\omega^2 - \omega^2 - \omega^2 \int_0^\infty dv v^2 \rho(v)^2\right]^{-1},
\]
(73)
where we have used Eq. (56). Thus, the $\omega$ dependence of $\xi(\omega)$ is not consistent with the causality requirements usually imposed on response functions, i.e., with Kramers-Kronig relations. This is because we have calculated a Feynman propagator and not a retarded Green’s function. The dielectric relations. This is because we have calculated a Feynman propagator in complete analogy with the formula for the free-space propagator Eq. (70). Here, $k_{cd} = \sqrt{\xi(\omega)\omega^2 - q^2}$ is the $z$ component of the complex wave vector in the medium with an always positive imaginary part.

We may now proceed by substituting Eq. (66) into Eq. (67),
\[
D_{ik}(z,z') = D_{ik}^{(e)}(z - z') - \frac{K(\omega)}{\epsilon_0} \int_0^\infty dz_1 D_{ij}^{(e)}(z - z_1)D_{jkl}(z_1 - z')
\]
\[
- \frac{K^2(\omega)}{\epsilon_0^2} \int_0^\infty dz_1 \int_0^\infty dz_2 D_{ij}^{(e)}(z - z_1)D_{jkl}(z_1 - z_2)\int_0^\infty d\xi \rho(\xi)^2 \mathbf{e}^{ik\xi z - i\xi q_{ik}}
\]
(75)
and focusing our attention on the solution of the cases $z < 0$ and $z' > 0$, i.e., when the source is located outside the dielectric and the observation point is inside the material. The solution for the case $z, z' > 0$ can then be obtained by applying the integral equation (66). The advantage of introducing Eq. (75) is that it facilitates iteration as it turns out that, when $D_{ik}$ on the right-hand side is replaced by $D_{ik}^{(e)}$, the action of the double-integral operator in the last term reduces to a matrix multiplication,
\[
\int_0^\infty dz_1 \int_0^\infty dz_2 D_{ij}^{(e)}(z - z_1)D_{jkl}(z_1 - z_2)D_{ik}^{(e)}(z_2 - z') = C_{ij} D_{ik}^{(e)}(z - z'),
\]
(76)
with the matrix $C_{ij}$ independent of $z$ and $z'$. In order to efficiently verify and make use of assertion (76), let us point out some useful facts. First, we recall that
\[
q^2 \delta_{ik} - q_i q_k = \omega^2 \left[\mathbf{e}_{i}^{\text{TE}}(q)\mathbf{e}_{k}^{\text{TE}}(q) + \mathbf{e}_{i}^{\text{TM}}(q)\mathbf{e}_{k}^{\text{TM}}(q)\right] - \rho(q)^2 - \omega^2 \int_0^\infty dv v^2 \rho(v)^2 = \frac{1}{\epsilon_0} \left[\omega^2 - \omega^2 - \omega^2 \int_0^\infty dv v^2 \rho(v)^2\right]^{-1},
\]
(77)
where $q = (q_x, k_z)$ is the wave vector in vacuum and we have introduced the polarization vectors,
\[
\mathbf{e}_{i}^{\text{TE}}(q_i) = \frac{1}{|q_i|} (-q_y, q_x, 0),
\]
\[
\mathbf{e}_{i}^{\text{TM}}(q_i, k_z) = \frac{1}{|q_i|} (q_z, k_z, q_x, 0),
\]
(78)
\[
\mathbf{e}_{i}^{\text{TM}}(q_{i,z}, k_{z,d}) = \frac{1}{|q_i|} \sqrt{\omega^2 - q_{z,d}^2} (q_k, k_{z,d}, q_{i,z}, 0).
\]
(77)
We have listed $\mathbf{e}_{i}^{\text{TM}}(q_i, k_z)$ explicitly to point out the additional factor of $\xi^{-2/3}(\omega)$ in its normalization. In the following, we will suppress the insignificant dependence of the polarization vectors on $q_i$. Relation (77) is simply a statement of the completeness property of the polarization vectors (78), but it allows us to write
\[
(\nabla \cdot \nabla_k - \delta_{jk} \nabla^2) e_{ik}(r_i - r_j) = i\omega^2 e_{ik}(r_i - r_j) e^{ik_{(z)z'}} + \frac{e_{ik}(r_i - r_j)}{2k_z} e^{ik_{(z')z}} e^{ik_{(z)z'}} + \frac{e_{i}(r_i - r_j)}{2k_z} e^{ik_{(z)z'}} e^{ik_{(z')z}}
\]
(77)
so that the partial Fourier transform of the free-space propagator (70) may be written as
\[
D_{ij}^{(0)}(z - z') = -i \frac{\epsilon_0 \omega^2}{2k_z} \sum_{\lambda} \left[\frac{\epsilon_{(z)z'}^{(0)}(k_{cd})\epsilon_{(z')z'}^{(0)}(k_{cd})e^{ik_{(z)z'}}}{\epsilon_{z'}^{(0)}(k_{cd})\epsilon_{z'}^{(0)}(k_{cd})e^{ik_{(z')z}}}\right]
\]
(77)
We emphasize that the above representation of the free-space propagator is not valid at the point $z = z'$ where the $z$ derivatives in Eq. (70), acting on $e^{ik_{(z)z'}}$, would produce additional terms proportional to a $\delta$ function. Similarly, we have
\[
D_{ij}^{(e)}(z - z') = -i \frac{\epsilon_0 \omega^2}{2k_{cd}} \sum_{\lambda} \left[\frac{\epsilon_{(z)z'}^{(e)}(k_{cd})\epsilon_{(z')z'}^{(e)}(k_{cd})e^{ik_{(z)z'}}}{\epsilon_{z'}^{(e)}(k_{cd})\epsilon_{z'}^{(e)}(k_{cd})e^{ik_{(z')z}}}\right]
\]
(77)
Equations (80) and (81) show that the free-space and bulk-medium propagators can be split into separate contributions from the transverse electric and transverse magnetic polarizations,
\[
D_{ij}^{(e)}(z - z') = \sum_{\lambda} D_{ij}^{(e,\lambda)}(z - z').
\]
(82)
Most of the further calculations are very much simplified if one takes into account that scalar products of polarization vectors with different $z$ components are diagonal in the polarization indices, i.e., we have
\[
\epsilon_{i}^{(e)}(q_i, z)\epsilon_{j}^{(e)}(q_j, z) = f^{e}(q_z, p_z) \delta_{i,j},
\]
(83)
The function $f$ is equal to 1 for the TE mode, and for the TM mode, it reads

$$f^{\text{TM}}(q_z, p_z) = \frac{q_z p_z + q_z^2}{\sqrt{q_z^2 + p_z^2}}. \tag{84}$$

This is very useful because it shows that not only a single propagator, as in Eq. (82), but also a product of propagators can always be split into separate contributions from the transverse electric and transverse magnetic modes, i.e., we can always write

$$\cdots D_{ij}(z - z_1) D_{kl}(z_1 - z_2) D_{ik}(z_2 - z') \cdots \equiv \sum_\lambda D_{\lambda,ij}(z - z_1) D_{\lambda,kl}(z_1 - z_2) D_{\lambda,ik}(z_2 - z') \cdots$$

This is true for an arbitrary number of propagators.

We can now proceed to verifying Eq. (76). First, we note that the arguments of all three propagators entering Eq. (76) have a definite sign. Indeed, we have

$$z - z_1 < 0, \quad z_1 - z_2 > 0, \quad z_2 - z' < 0. \tag{85}$$

Thus, from Eqs. (80) and (81), it follows that the propagators entering the integral in Eq. (76) are given by

$$D^{(0)}_{ij}(z - z') = - \frac{i \epsilon_0 \omega^2}{2k_z} e^{ik_z(z - z')} \sum_\lambda e^{ik_z(z)} e^{ik_z(z')},$$

$$D^{(e)}_{ij}(z - z') = - \frac{i \epsilon_0 \xi^2(\omega) \omega^2}{2k_{cd}} e^{-ik_{cd}z} \sum_\lambda e^{ik_z(z)} e^{ik_z(z')}, \tag{86}$$

With this, we can evaluate the integrals in Eq. (76) and find that

$$\int_0^\infty d z_1 \int_0^\infty d z_2 D^{(e)}_{ij}(z - z_1) D^{(0)}_{ij}(z_1 - z_2) D^{(e)}_{ik}(z_2 - z') = \frac{\epsilon_0^4}{K^2(\omega)} \sum_\lambda \frac{r^2_\lambda}{1 - r^2_\lambda} D^{(e)}_{\lambda,ik}(z - z'), \tag{87}$$

where we have used Eq. (73). Here, $r_\lambda$ is the Fresnel coefficient for reflection from a half-space. Since all the Fresnel coefficients for reflection and transmission at a half-space are needed later on, we list them here

$$r_{\text{TE}} = \frac{k_z - k_{cd}}{k_z + k_{cd}}, \quad r_{\text{TM}} = - \frac{\xi(\omega) k_z - k_{cd}}{\xi(\omega) k_z + k_{cd}},$$

$$t_{\text{TE}} = \frac{2k_z}{k_z + k_{cd}}, \quad t_{\text{TM}} = - \frac{2 \sqrt{\xi(\omega) k_z}}{\xi(\omega) k_z + k_{cd}}. \tag{88}$$

The significance of Eq. (87) is that it allows us to iterate the integral equation (75) along the lines of Eqs. (63) and (64). The iterative process is straightforward and, thanks to relation (87), leads to two separate geometric series for the two polarizations,

$$D_{\lambda,ij}(z - z') = \left[D^{(e)}_{\lambda,ij}(z - z') - \frac{K(\omega)}{\epsilon_0^2} \int_0^\infty d z_1 D^{(e)}_{\lambda,ij}(z - z_1) D^{(0)}_{\lambda,ik}(z_1 - z') \right] \times \left[1 - \frac{r^2_\lambda}{1 - r^2_\lambda} + \frac{r^2_\lambda}{1 - r^2_\lambda} \cdots \right]. \tag{89}$$

These geometric series can easily be summed up for all orders to give the exact photon propagator for the cases $z < 0, z' > 0$

In order to cast the result into a familiar form, we explicitly evaluate the integral in the second line, which requires some care. The integral that needs to be evaluated is

$$I^i_{\lambda}(z, z') = \frac{K(\omega)}{\epsilon_0^2} \int_0^\infty d z_1 D^{(e)}_{\lambda,ij}(z - z_1) D^{(0)}_{\lambda,ik}(z_1 - z'). \tag{90}$$

Here, the argument of $D^{(e)}$ is always negative, $z - z_1 < 0$, whereas, the sign of $z_1 - z'$ can be both positive and negative. Therefore, we need to take into account that the propagator (70) is discontinuous at $z_1 = z'$ and contains singular terms proportional to $\delta(z_1 - z')$. In order to correctly evaluate the integral (90), we represent the differential operator in Eq. (70) using the polarization vectors written out in terms of derivatives. Using the completeness relation of the transverse polarization vectors, we may symbolically write

$$\nabla_i \nabla_k - \delta_{ik} \nabla^2 = -\nabla^2 \sum_\lambda e^\lambda_i(\nabla) e^\lambda_k(\nabla). \tag{91}$$

With this, the propagators entering the integral (90) are given by

$$D^{(0)}_{\lambda,ij}(z_1 - z') = - \frac{i \epsilon_0}{2k_z} \left(q^\lambda_\parallel - \nabla^\lambda_z\right) e^\lambda_i(-\nabla^\lambda_z) e^\lambda_j(-\nabla^\lambda_z) e^{ik_z(z_1 - z')}, \tag{92}$$

$$D^{(e)}_{\lambda,ij}(z - z_1) = - \frac{i \epsilon_0 \xi^2(\omega) \omega^2}{2k_{cd}} e^{ik_{cd}(z_1 - z')} \sum_\lambda e^\lambda_i(-k_{cd}) e^\lambda_j(-k_{cd}) e^{-ik_{cd}(z_1 - z')} \tag{93}$$

Note that, in $D^{(0)}$, we have changed the $z$ derivatives to act on $z'$ rather than on $z$ so that they could be pulled outside the integral in Eq. (90). Now, it is straightforward to demonstrate that the integral (90) is given by

$$I^i_{\lambda}(z, z') = D^{(e)}_{\lambda,ij}(z - z') - \frac{i \epsilon_0 \xi(\omega) \omega^2}{2k_{cd}} \sum_\lambda e^\lambda_i(-k_{cd}) e^\lambda_j(-k_{cd}) e^{-ik_{cd}(z + k_z z')}, \tag{94}$$

whose first term exactly cancels the bulk dielectric propagator in the first line of Eq. (89). The remaining term yields the final result,

$$D_{ij}(z, z') = - \frac{i \epsilon_0 \xi^2(\omega) \omega^2}{2k_{cd}} \sum_\lambda \left[\xi(\omega) e^\lambda_i(-k_{cd}) e^\lambda_j(-k_{cd}) t_k e^{-ik_{cd}(z + k_z z')}, \tag{95}$$

with the transmission coefficient as given in Eq. (88). This formula describes the vacuum-dielectric transmission, i.e., it is valid for $z < 0, z' > 0$. It is a straightforward calculation to plug Eq. (95) into Eq. (66) and to obtain the photon propagator for the case $z, z' > 0$. In the region $z' > 0$, the final result for the dressed photon propagator Fourier-transformed back to
coordinate space, may be written as

\[
D_{ij}(r, r'; \omega) = \theta(z)D_{ij}^{(0)}(r - r'; \omega) - \frac{i\epsilon_0}{(2\pi)^2} \sum_{q} \int d^2 q \frac{\omega^2}{2k_z} e^{iq_z(r_z-r'_z)} \times \left\{ \begin{array}{l}
\theta(-z)\left[ \xi_{\omega}(q_{q_1} - k_{zq_1}, -k_{zq_1}) \epsilon_{\omega}(q_{q_1} - k_{zq_1}, -k_{zq_1}) e^{-i(k_{zq_1} + q_z)z}
+ \theta(z)\left[ \epsilon_{\omega}(q_{q_1}, -k_{zq_1}) \epsilon_{\omega}(q_{q_1}, -k_{zq_1}) e^{i(k_{zq_1} + q_z)z} \right]\end{array} \right\}.
\]  

In the calculations of the energy-level shifts of an atom placed outside an absorbing dielectric material to be discussed in the following section, we will need the propagator for the case \( z, z' > 0 \). In that case, Eq. (96) shows that the propagator splits into a free-space part \( D_{ij}^{(0)}(r - r'; \omega) \), which is not interesting as it just yields the standard (position-independent) Lamb shift, and a correction due to reflection at the boundary, which we call \( D_{ij}^{(\alpha)}(r, r'; \omega) \) and which gives rise to the position-dependent Casimir-Polder shift. As we treat the atom-field interaction in the dipole approximation, we are going to need the reflected part of the propagator \( D_{ij}^{(\alpha)}(r, r'; \omega) \) evaluated at equal arguments \( r = r' = R \), where \( R = (0, 0, Z) \) is the position of the atom. In that case, it simplifies considerably and can be written in the form

\[
D^{(\alpha)}(Z; \omega) = -\frac{i\epsilon_0}{8\pi} \int_0^\infty dq_1 \frac{q_1}{k_z} e^{2i(k_z - q_z)Z} \times \left( \begin{array}{ccc}
\omega^2 r_{TE} - k_z^2 r_{TM} & 0 & 0 \\
0 & \omega^2 r_{TE} - k_z^2 r_{TM} & 0 \\
0 & 0 & 2q_1^2 r_{TM}^2 \end{array} \right),
\]

with \( k_z = \sqrt{\omega^2 - q_1^2} + i\eta \) as before. Note that we have gone to polar coordinates \( q_z = q_1 \cos \phi, q_r = q_1 \sin \phi \) where the azimuthal integration annihilated the off-diagonal elements of the equal-argument propagator \( D_{ij}^{(\alpha)}(r, r; \omega) \).

As a final remark, we would like to comment on the convergence of the series in Eq. (89). It clearly converges provided

\[
\left| \frac{r_z}{1 - r_z^2} \right| < 1.
\]

However, there does not seem to be a physical significance to this condition. That the result for the propagator can be extended by analytic continuation to wave vectors not satisfying the condition (98) can be shown by solving the corresponding boundary-value problem, the procedure for which we sketch in Appendix A.

V. ATOMIC PROPAGATOR AND ELECTRON SELF-ENERGY

In order to investigate the perturbative expansion of the atomic propagator (28), we use the expansion in terms of atomic eigenstates, Eq. (24), and then work with the atomic propagator in that basis. In analogy to Eqs. (30) and (31) for the unperturbed propagator, we obtain

\[
G_{ii}(t, t') = \sum_{n=0}^{\infty} \left( -\frac{i}{\hbar} \right)^n \int dt_1 \cdots \int dt_n \langle \Omega | \prod_{n=1}^{n} T \{ c_n(t) c_n(t') H_{A-EM}(t_n) \} | \Omega_{\text{conn}} \rangle.
\]

By using Wick’s theorem to evaluate the ground-state expectation value of the time-ordered product of operators, one easily sees that the zeroth-order term is a propagator for the noninteracting system and that the first-order correction vanishes because it is not possible to contract all of the operators. Therefore, the lowest-order nonvanishing perturbative contributions come from terms of order \( e^2 \). Diagrams of this order have two vertices and, therefore, include a pair of disconnected tadpole diagrams, which are irrelevant as they go away in the process of normalization, and the physically important self-energy diagram which contains all the information about the energy-level shifts and decay rates.

For calculating the perturbatively corrected dressed atomic propagator, it is, in fact, convenient to perform a partial summation and to consider the following series of diagrams:

\[
= - + \quad \text{with the thick solid line represents the dressed atomic propagator.}
\]

where the location of the poles of such-constructed propagator is much more straightforward to work out than for the propagator with strictly only one-loop corrections. The above is a graphical representation of the Dyson equation, which expressed analytically reads

\[
G_{ii}(t, t') = G_{ii}^{(0)}(t, t') + \frac{i\hbar}{\epsilon_0^2} \sum_{k, l, m, n} \mu_{ml}^{ik} \mu_{ln}^{ik} \int_{-\infty}^\infty dt_1 \int_{-\infty}^\infty dt_2 \times G_{nn}^{(0)}(t, t_1) G_{nn}^{(0)}(t_1, t_2) D_{ij}(R, R, t_1, t_2) G_{ij}(t_2, t').
\]

We note that, here and in the following, the indices \( k \) and \( l \) label just Cartesian components, but the sum over \( m \) is a sum over intermediate atomic eigenstates, and \( i \) is the atomic state whose energy shift we are seeking to determine. To this end, we Fourier transform (100) with respect to \( t - t' \) along the lines of Eq. (34) and find the dressed atomic propagator,

\[
G_{ii}(E) = \int_{-\infty}^\infty dt(t - t') e^{i(t-t')E/\hbar} G_{ii}(t - t') = \frac{1}{E - E_i + i\eta - \Sigma_{ii}(E)}.
\]
with the self-energy insertion,
\[ \Sigma_{ii}(\mathcal{E}) = \frac{i\hbar}{2\pi\varepsilon_0} \sum_{k,l,m} \mu_{mi}^k \mu_{lm}^k \int_{-\infty}^{\infty} d\omega \frac{D_{kl}(\mathbf{R}, \mathbf{R}; \omega)}{\mathcal{E} - \hbar \omega - E_m + i\eta}, \]  
(102)

The self-energy insertion (102) contains the dressed photon propagator which, in the case of an atom outside a dielectric half-space, comprises: (i) the free-space photon propagator \( D_{kl}^{(0)}(\mathbf{R}, \mathbf{R}; \omega) \), which gives rise to the position-independent Lamb shift, and (ii) the reflected part \( D_{kl}^{(r)}(\mathbf{R}, \mathbf{R}; \omega) \), which yields the position-dependent Casimir-Polder shift. Thus, the shift in the atomic energy levels, given by the poles of Eq. (101), can be written
\[ \mathcal{E} - E_i = \Sigma_{ii}^{(0)}(\mathcal{E}) + \Sigma_{ii}^{(r)}(\mathcal{E}). \]  
(103)

As we want to work out changes in the energy levels already corrected for the coupling between the atom and the free-space electromagnetic fields, we renormalize the energy-level shift by subtracting the self-energy associated with the free-space electromagnetic field \( \Sigma_{ii}^{(0)}(\mathcal{E}) \) and consider
\[ \Delta \mathcal{E}_i^{\text{ren}} = \mathcal{E} - \bar{E}_i = \Sigma_{ii}^{(r)}(\mathcal{E}). \]  
(104)

We use the symbol \( \bar{E}_i \equiv E_i + \Sigma_{ii}^{(0)}(\mathcal{E}) \) to represent the atomic energy levels already corrected for the free-space Lamb shift and decay rates.

\textit{Prima facie} it may seem difficult to extract the energy shift from Eq. (104) because it is an implicit equation whose right-hand side also depends on \( \mathcal{E} \). However, if the energy shift we calculate is small compared to the difference in energy between the state under consideration and its nearest dipole-connected neighbors (which it needs to anyhow for perturbation theory to be applicable), then the shift can be extracted from Eq. (104) by a single iteration leading to
\[ \Delta \mathcal{E}_i^{\text{ren}} \approx \Sigma_{ii}^{(r)}(\bar{E}_i). \]  
(105)

where we have abbreviated \( \omega_{mi} = \omega_m - \omega \). The \( \omega \) integral in Eq. (105) can be restricted to the positive real axis by writing
\[ \frac{1}{\omega + \omega_{mi} - i\eta} = \frac{\omega - \omega_{mi}}{\omega^2 - (\omega_{mi} - i\eta)^2}, \]  
(106)

and noting that \( D_{kl}^{(r)}(\mathbf{R}, \mathbf{R}; \omega) \) is even in \( \omega \) (see Sec. IV B and Appendix B). Then, the term proportional to \( \omega \) is odd and vanishes when integrated over the real \( \omega \) axis. As the photon propagator is analytic in the first quadrant of the complex \( \omega \) plane, it is permissible, provided \( \omega_{mi} > 0 \), to rotate the contour of \( \omega \) integration by \( \pi/2 \), i.e., \( \omega \to i\omega \). This applies when one considers an atom in its ground state.

However, for an excited state \( i \) of the atom, one has \( \omega_{mi} < 0 \), which means that there will be poles in the first quadrant of the \( \omega \) plane due to the denominator in Eq. (105). We would also like to remark that the Fresnel reflection coefficients have poles in the complex plane at the location of trapped electromagnetic modes, which is not an issue in the case of a dielectric half-space but arises, e.g., for a dielectric slab [29] and other systems capable of waveguiding [30].

We recall that \( D_{kl}^{(r)}(\mathbf{R}, \mathbf{R}; \omega) \) is diagonal, cf. Eq. (97) and write down the final result for the energy shift in the form
\[ \Delta \mathcal{E}_i^{\text{ren}} = \Delta \mathcal{E}_i + \Delta \mathcal{E}_i^r. \]  
(107)

with \( \Delta \mathcal{E}_i \) and \( \Delta \mathcal{E}_i^r \) given by
\[ \Delta \mathcal{E}_i = \frac{1}{\pi\varepsilon_0} \sum_{k, m} |\mu_{mi}^k|^2 \int_{0}^{\infty} d\omega \frac{\omega_{mi}}{\omega^2 + \omega_{mi}^2} D_{kl}^{(r)}(\mathbf{R}, \mathbf{R}; i\omega), \]  
(108)

\[ \Delta \mathcal{E}_i^r = \frac{1}{\varepsilon_0} \sum_{k, m} |\mu_{mi}^k|^2 D_{kl}^{(r)}(\mathbf{R}, \mathbf{R}; |\omega_{mi}|) \theta(-\omega_{mi}). \]  
(109)

where \( |\mu_{mi}^k| \equiv |\langle m| k \rangle | \) are the matrix elements of the \( k \)th component of the electric dipole moment operator. The quantity \( \Delta \mathcal{E}_i^r \) is the contribution to the self-energy that originates from the poles in Eq. (105) that arise for an excited state \( i \) for which \( \omega_{mi} < 0 \). Expressions equivalent to Eqs. (108) and (109) have been derived before by different methods, e.g., by linear-response theory [31,32] or later by the noise-current approach to phenomenological QED [33].

The shift \( \Delta \mathcal{E}_i \) is real because it is a convolution of atom and field susceptibilities, which are real at complex frequencies [31]. However, \( \Delta \mathcal{E}_i^r \) is complex; its imaginary part modifies the decay rates of excited states. To summarize, we have
\[ \Delta E_i = \Delta \mathcal{E}_i + \text{Re}(\Delta \mathcal{E}_i^r), \]
\[ \Delta \Gamma_i = -\frac{2}{\hbar} \text{Im}(\Delta \mathcal{E}_i^r), \]  
(110)

where \( \Delta E_i \) are the renormalized energy-level shifts and \( \Delta \Gamma_i \) are the changes in decay rates.

VI. ENERGY-LEVEL SHIFTS NEAR A HALF-SPACE

A. Ground state

Substituting the photon propagator (97) into Eq. (108), we find that the energy shift of the atomic ground state \( |g\rangle \) is given by
\[ \Delta E_g = -\frac{1}{8\pi^2\varepsilon_0} \sum_{m} \int_{0}^{\infty} d\omega_{mg} \int_{0}^{\infty} d\omega_{mg} \int_{0}^{\infty} d\omega_{mg} e^{-\sqrt{\frac{\omega_{mg}^2}{2} + \omega^2}} \times \left[ (q_{\|}^2 + \omega_{mg}^2)^2 - \omega_{mg}^2 \right] |\mu_{mg}^1|^2 + 2q_{\|}^2 |\mu_{mg}^1|^2 |\mu_{mg}^2|^2, \]  
(111)

where we have used the notation \( |\mu_{mg}^1|^2 = |\mu_{mg}^1|^2 + |\mu_{mg}^2|^2 \) and \( |\mu_{mg}^2|^2 = |\mu_{mg}^2|^2 \). The reflection coefficients are as defined in Eq. (88). In terms of the new variables, they read
\[ q_{\|}^2 = \sqrt{\omega_{mg}^2 + q_{\|}^2} = \sqrt{\frac{\omega^2 + \omega_{mg}^2}{\omega^2 + \omega_{mg}^2}}, \]
\[ q_{\perp}^2 = \sqrt{\frac{\omega_{mg}^2 + q_{\perp}^2}{\omega_{mg}^2 + q_{\perp}^2}} = \sqrt{\frac{\omega^2 + \omega_{mg}^2}{\omega^2 + \omega_{mg}^2}}, \]  
(112)

Note that we have replaced \( \xi(\omega) \) by \( \epsilon(\omega) \) in Eq. (112) compared to Eq. (88) because, for the relevant frequencies, both functions coincide (see Sec. IV B and Appendix B).
If we now introduce polar coordinates according to 
\( \omega = \omega_{mg} \cos \phi, \quad q_{\perp} = \omega_{mg} \sin \phi \) and then write \( y = \cos \phi \), we obtain the, perhaps, most useful expression for the ground-state shift, especially for numerical analysis and for investigating the effects of retardation,

\[
\Delta E_g = -\frac{1}{8\pi^2\epsilon_0} \sum_m \int_0^\infty dx x^3 \int_0^1 dy \frac{\alpha_m^3}{1 + x^2 y^2} e^{-2\omega_{mg} Z x} \times (|\tilde{p}_{TM} - \gamma^2 \tilde{p}_{TE}|\mu_{mg}^\parallel |^2 + 2(1 - \gamma^2)\tilde{p}_{TM}^\parallel |\mu_{mg}^\parallel |^2).
\]

(113)

The result in Eq. (113) formally takes the same form as the results obtained in calculations involving only nondispersive dielectrics (see, e.g., Ref. [19]), the only difference being the reflection coefficients that now, through the dielectric constant, depend on the product \( xy \) of the integration variables, which is the photon frequency in units of \( \omega_{mg} \).

\[
p_{TE} = \frac{1 + \sqrt{1/\epsilon((i\omega_{mg}xy) - 1)^2 + 1}}{1 + \sqrt{1/\epsilon((i\omega_{mg}xy) - 1)^2 + 1}},
\]

(114)

\[
p_{TM} = \frac{\epsilon((i\omega_{mg}xy) - 1)^2 + 1}{\epsilon((i\omega_{mg}xy) + 1)^2 + 1}.
\]

Equation (113) is suitable for numerical analysis but does not give immediate insight into the dependence of the energy shift as a function of the distance from the surface. It is, therefore, instructive to consider some of its limiting cases.

As has been spelled out, e.g., in Ref. [19], the dimensionless parameter that plays a decisive role in the characteristics of the Casimir-Polder interaction is given by the combination \( 2\omega_{mg} Z \), which is the ratio of two time scales: (i) the typical time \( 2Z/c \) needed by a virtual photon to make a round trip between the atom and the surface and (ii) the typical time scale \( \omega_{mg}^{-1} \), at which, the atomic system evolves. While Eq. (113) includes a sum over atomic states \( |n\rangle \), in reality, contributions to the state are dominated by the state which is connected to the ground state by the strongest dipole transition. We will call the frequency \( \omega_{mg} \), that pertains to this strongest transition the “typical transition frequency,” and it is this number that enters the retardation criterion parameter. Roughly speaking, if \( 2\omega_{mg} Z \ll 1 \), we are in the so-called nonretarded regime when the time needed by the photon to travel between the dielectric and the atom is negligibly small compared to the typical atomic time scale. Then, the interaction can safely be approximated as instantaneous, and our result should reduce to that calculated by Barton [34], who considers only the Coulomb interaction of an atom with surface polaritons. In the opposite case \( 2\omega_{mg} Z \gg 1 \), the interaction becomes retarded, i.e., by the time the photon has completed a round trip, the atomic state has changed significantly. In that case, for reasons that are not obvious but will become apparent later, the interaction depends only on static polarizabilities, i.e., the polarizabilities evaluated at zero frequency. The diagonal polarizability of the spherically symmetric atom is then

\[
\alpha_{s\parallel}(0) = \sum_j \frac{2\omega_{ji}}{\omega_j^2 - \omega^2} \left| \mu_{ji} \right|^2.
\]

and the susceptibility of the dielectric becomes

\[
\epsilon(0) = \frac{\epsilon_0}{\epsilon_0 - \omega^2 - 2i\gamma_0} = 1 + \frac{\alpha_0^2}{\omega_0^2},
\]

(116)

as explained in Appendix B.

1. Nonretarded limit

In order to take the nonretarded limit of the energy shift, it is best to start from Eq. (111). After changing variables from \( \omega \) to \( s \), we take the limit \( 2\omega_{mg} Z \rightarrow 0 \), which we may do because the line \( s = \infty \) does not contribute to the integral, and approximate

\[
q_{\perp}^2 + \omega^2 \approx q_{\perp}^2 [1 + (2\omega_{mg} Z)^2 s^2] \approx q_{\perp}^2 + \omega^2 \rightarrow q_{\perp}^2 [1 + (2\omega_{mg} Z)^2 s^2] \approx q_{\perp}^2.
\]

This significantly simplifies Eq. (111). The \( q_{\perp} \) integral becomes elementary, and the final result reads

\[
\Delta E_g \approx -\frac{1}{32\pi^2\epsilon_0 Z^2} \sum_m \int_0^\infty d\omega \omega_{mg} \epsilon(i\omega) - 1 - \omega^2 + \omega_{mg} \epsilon(i\omega) + 1 \times (|\mu_{mg}^\parallel|^2 + 2|\mu_{mg}^\parallel|^2).
\]

(117)

We observe the expected \( Z^{-3} \) behavior of the energy-level shift in the nonretarded or “van der Waals” regime. The exact energy-level shift for all distances is plotted in Fig. 3; as the plot shows the energy shift multiplied by \( Z^2/\omega_{mg} \), the linear rise in the graph at small arguments represents the \( Z^{-3} \) behavior derived above.

In order to see that the result in Eq. (117) is equivalent to the slightly more awkward principal-value integral given in Eq. (7.14) of Ref. [34] or Eq. (13) of Ref. [35], one needs to rewrite

\[
\frac{\omega_{mg}}{\omega^2 + \omega_{mg}^2} = \frac{1}{2} \left( \frac{1}{\omega_{mg} - i\omega} + \frac{1}{\omega_{mg} + i\omega} \right),
\]

and to re-rotate the contour from \( \omega \) to \( i\omega \) in the first and to \(-i\omega \) in the second summand. Equation (117) also confirms the results derived on the basis of the phenomenological noise-current approach to quantum electrodynamics with dielectric media, see, e.g., Ref. [33].

2. Retarded limit

In order to work out an approximation for the energy shift when retardation is dominant, it is convenient to start with Eq. (113) where the decisive parameter \( 2\omega_{mg} Z \) is present in the exponential, which, in the limit \( 2\omega_{mg} Z \rightarrow \infty \), strongly damps the integrand. Then, the main contributions to the integral come from the neighborhood of \( x = 0^+ \), and one can obtain an asymptotic expansion of the integral by expanding the integrand in a Taylor series around this point. A straightforward calculation gives

\[
\Delta E_{g,e} \approx -\frac{3}{64\pi^2\epsilon_0} \sum_m \sum_{s = 1, L} \left( \frac{e^2}{\omega_0^2 - \omega^2} - 4\gamma \frac{e^2}{\omega_0^2 - \omega^2} \right) \frac{|g| |\mu^\parallel|^2}{\omega_{mg}^3}.
\]

(118)

where, in the parentheses, we have neglected terms of order \( \omega_{mg}^{-5} Z^{-6} \) and higher. The coefficients \( c_{s,L} \) are given below, depending only on the static dielectric constant of the material.
The fact that, to leading order, the Casimir-Polder force depends only on the static polarizability of the atom, Eq. (115), is well known [31]. Therefore, the leading-order \( Z^{-4} \) term in Eq. (118) is identical to the retarded limit of the energy shift in a ground-state atom interacting with a nonabsorptive dielectric half-space described by a static refractive index \( n(0) \equiv n = (1 + \omega_T^2/\omega_r^2) \), which has been derived previously [19]. We just quote the results for the coefficients \( c_4, c_5 \) from Ref. [19],

\[
\begin{align*}
\frac{c_4}{n^2 - 1} & = -\frac{1}{3} \left( \frac{2}{3} n^2 + n - \frac{8}{3} \right) + \frac{2n^4}{(n^2 - 1)\sqrt{n^2 + 1}} \ln \left( \frac{\sqrt{n^2 + 1} + 1}{n[\sqrt{n^2 + 1} + n]} \right) + \frac{2n^4 - 2n^2 - 1}{(n^2 - 1)^{3/2}} \ln(\sqrt{n^2 + 1} + n), \\
\frac{c_5}{n^2 - 1} & = \frac{1}{3} \left( \frac{4n^4 - 2n^3 - 4}{3} n^2 + 4 \right) - \frac{4n^6}{(n^2 - 1)\sqrt{n^2 + 1}} \ln \left( \frac{\sqrt{n^2 + 1} + 1}{n[\sqrt{n^2 + 1} + n]} \right) - \frac{2n^7(2n^4 - 2n^2 + 1)}{(n^2 - 1)^{3/2}} \ln(\sqrt{n^2 - 1} + n).
\end{align*}
\]

In other words, to leading order, in the retarded limit, absorption makes no difference, and only static polarizabilities of both the dielectric and the atom matter. This is because the photon wavelengths that matter the most in the atom-wall interaction are of the order of the distance between the atom and the surface of the dielectric and longer. Thus, for an atom in the so-called far zone, only long wavelengths of the electromagnetic radiation come into play, which means low frequencies. This is illustrated by the plot of the exact energy-level shift in Fig. 3; for large arguments, all lines tend to the same value, and absorption just determines how quickly.

Now, we turn our attention to the next term in the asymptotic expansion, which is proportional to \( Z^{-5} \). This is the first term that contains information about corrections to the energy shift due to absorption in the retarded regime. Apart from the factor \( 4\gamma/\omega_r^2 \), the dimensionless coefficients \( c_6, c_7 \) depend again only on the static refractive index \( n = \sqrt{1 + \omega_T^2/\omega_r^2} \) and are given by

\[
\begin{align*}
\frac{c_6^\perp}{3(n - 1)(n + 1)^2(n^2 + 1)} & = \left\{ 4n^6 - 3n^5 - 11n^4 + 4n^3 + 2n^2 - 5n + 7 \right. \\
& \left. - 6n^2(n^5 + n^4 - n^3 - n^2 - 2n - 2) \ln \left[ n \left( \frac{n + 1}{n^2 + 1} \right) \right] \right\}, \\
\frac{c_7^\parallel}{3(n - 1)(n + 1)^2(n^2 + 1)} & = \frac{4}{(n - 1)(n + 1)^2(n^2 + 1)} \times \left\{ -6n^8 + 3n^7 + 10n^6 - 5n^5 + 3n^4 - n^3 \\
& - 6n^2 + n + 1 + 3n^3(2n^5 + 2n^4 - n^3 - n^2 - 3n - 3) \ln \left[ n \left( \frac{n + 1}{n^2 + 1} \right) \right] \right\}.
\end{align*}
\]

We provide plots of these functions in Fig. 2 from where a quick estimate of the value of these coefficients can be obtained. Since both \( c_6^\perp \) and \( c_7^\parallel \) are positive, we see that absorption reduces the magnitude of the ground-state energy shift by an amount that is proportional to the damping constant \( \gamma \), cf. Fig. 3. We also note that the correction goes with the inverse square of the absorption frequency \( \omega_r \) in the dielectric so that only absorption lines that lie at sufficiently low frequencies make a significant difference. This happens because the main contribution to the ground-state shift in the retarded limit comes from long wavelengths or, equivalently, small values of \( x \) (which is a scaled frequency). Therefore, the integral is not sensitive to any absorption peaks which lie at higher frequencies as there, the integrand is highly damped anyway, cf. Eq. (113).
The shift of an excited energy level gets contributions from both parts of \( \Delta \mathcal{E}_\text{res} \), Eqs. (108) and (109). The nonresidue contributions, Eq. (108), assume exactly the same form as the results of the previous section. Therefore, we will not analyze them again but will, instead, have a closer look at the additional contributions due to Eq. (109).

Plugging in the photon propagator, Eq. (97), we find that the energy shift in the excited state \(|i\rangle\) is given by the real part of the following expression:

\[
\Delta \mathcal{E}_i^* = -\frac{i}{8\pi\epsilon_0} \sum_{m < i} \int_1^\infty dk_z e^{i\omega_m|Z_k|} \frac{d\eta}{\omega_m^2 - q_f^2 + i\eta} \left[|\omega_m|^2 r_{TM}^{\text{TE}} + (\omega_m^2 - q_f^2) r_{TM}^{\text{TM}}\right]|\mu_{m}^\parallel|^2 + 2q_f^2 |r_{TM}^{\text{TM}}|\mu_{m}^\bot^2. \tag{120}
\]

Here, \( r_{TM} \) are the reflection coefficients of Eq. (88), evaluated at the atomic transition frequencies \( \omega = |\omega_m| \). Also, the restriction of the sum over atomic states to those lying below state \(|i\rangle\) should be noted. For the purposes of asymptotic analysis of \( \Delta \mathcal{E}_i^* \), we change the integration variable in Eq. (120) to \( k_z = \sqrt{\omega_m^2 - q_f^2}/|\omega_m| \) and get

\[
\Delta \mathcal{E}_i^* = -\frac{i}{8\pi\epsilon_0} \sum_{m < i} |\omega_m|^3 \int_1^\infty dk_z e^{2i|\omega_m|Z_k_z} \times \left[ (f_{TM}^{\text{TE}} - k_z^2 f_{TM}^{\text{TM}})|\mu_{m}^\parallel|^2 + 2(1 - k_z^2) f_{TM}^{\text{TM}}|\mu_{m}^\bot|^2 \right], \tag{121}
\]

where the contour of integration runs from \( k_z = 1 \) along the real axis to \( k_z = 0 \) and then up along the imaginary axis to \( k_z = i\infty \). The reflection coefficients, expressed as functions of \( k_z \), are

\[
\begin{align*}
f_{TM}^{\text{TE}}(k_z) &= \frac{k_z - \sqrt{\epsilon(|\omega_m| - 1) + k_z^2}}{k_z + \sqrt{\epsilon(|\omega_m| - 1) + k_z^2}} \\
f_{TM}^{\text{TM}}(k_z) &= \frac{\sqrt{\epsilon(|\omega_m| - 1) + k_z^2}}{\epsilon(|\omega_m|)k_z + \sqrt{\epsilon(|\omega_m| - 1) + k_z^2}}\tag{122}
\end{align*}
\]

We now go on to analyze \( \Delta \mathcal{E}_i^* \) in the nonretarded and retarded limits.

### 1. Nonretarded limit

In the nonretarded limit of Eq. (121), we have \( 2|\omega_m|Z \ll 1 \). It is expedient to split the integration in Eq. (121) in the following way:

\[
\int_1^\infty dk_z = \int_0^\infty d(i\epsilon_k) - \int_0^1 dk_z, \tag{123}
\]

and to note that, in the limit \( 2|\omega_m|Z \to 0 \), the second integral on the RHS contributes to the asymptotic series only terms that are proportional to non-negative powers of \( Z \) and can, therefore, be discarded. The remaining part is given by

\[
\Delta \mathcal{E}_i^{*\text{,1,ret}} = -\frac{1}{8\pi\epsilon_0} \sum_{m < i} |\omega_m|^3 \int_0^\infty dk_z e^{-2i|\omega_m|Z_k_z} \times \left[ (f_{TM}^{\text{TE}} + k_z^2 f_{TM}^{\text{TM}})|\mu_{m}^\parallel|^2 + 2(1 + k_z^2) f_{TM}^{\text{TM}}|\mu_{m}^\bot|^2 \right]. \tag{124}
\]

where \( f_{TM}^{\text{TM}} \) are the reflection coefficients of Eq. (122) evaluated at imaginary argument \( f_{TM}^{\text{TM}} = f_{TM}^{\text{TM}}(ik_z) \).

Scaling the integration variable according to \( x = 2|\omega_m|Z_k_z \) and approximating

\[
\sqrt{\epsilon(|\omega_m| - 1) - \frac{x^2}{2|\omega_m|Z^2}} \approx \frac{i}{2|\omega_m|Z}, \tag{125}
\]

we derive that, in the nonretarded limit, Eq. (120) becomes

\[
\Delta \mathcal{E}_i^{*\text{,nonret}} = -\frac{1}{32\pi\epsilon_0 Z^3} \sum_{m < i} \frac{|\omega_m|^2}{\epsilon(|\omega_m|)} \left[ |\mu_{m}^\parallel|^2 + 2|\mu_{m}^\bot|^2 \right]. \tag{126}
\]

To leading order, the residue contributions to the energy shift of the excited state \(|i\rangle\), cf. Eq. (109), are given by the real part of the above expression,

\[
\Delta \mathcal{E}_i^{*\text{,nonret}} = -\frac{1}{32\pi\epsilon_0 Z^3} \sum_{m < i} \frac{|\omega_m|^2}{\epsilon(|\omega_m|)} \left[ |\mu_{m}^\parallel|^2 + 2|\mu_{m}^\bot|^2 \right]. \tag{127}
\]

Thus, in the nonretarded regime, the residue contributions behave as \( Z^{-3} \) and, therefore, are on the same order as the nonresidue contributions, cf. Eq. (117). The result in Eq. (127) is, in fact, equivalent to the real part of Eq. (7.10), derived in Ref. [34].

### 2. Retarded limit

Now, we turn our attention to the asymptotic behavior of Eq. (121) in the retarded limit, i.e., when \( 2|\omega_m|Z \gg 1 \). It is again useful to split the integration in the same way as in Eq. (123), only that now, both integrals play an important role. The first contribution, the integral along \( k_z \in [0,i\infty) \), given in Eq. (124), can be tackled by use of Watson’s lemma [36]. Noting that for \( 2|\omega_m|Z \gg 1 \), the integrand is strongly damped, we separate off the exponential and expand the remaining part into Taylor series about \( k_z = 0 \). The resulting integrals are elementary, and we obtain, for the leading term,

\[
\Delta \mathcal{E}_i^{*\text{,1,ret}} = -\frac{1}{8\pi\epsilon_0} \sum_{m < i} |\omega_m|^3 \left[ |\mu_{m}^\parallel|^2 \right] \frac{2|\omega_m|^2}{2|\omega_m|Z} \left[ 1 - \frac{2i\epsilon(|\omega_m|)}{\epsilon(|\omega_m|) - 1/2|\omega_m|Z} \right]^2 |\mu_{m}^\bot|^2 \tag{128}
\]

Next, we deal with the integral on the interval \( k_z \in [0,1] \) which, unlike in the nonretarded case, cannot be discarded. However, its asymptotic expansion in inverse powers of \( Z \) is easily obtained by repeated integration by parts. Interestingly, the asymptotic series contain nonscillatory terms that exactly
cancel out the contributions given in Eq. (128). Altogether, we find that the leading-order and next-to-leading-order terms are
\[ \Delta \mathcal{E}_{i}^{\text{ret}} = \frac{1}{4\pi \epsilon_0} \sum_{m<i} \frac{\omega_{mi}}{|\omega_{mi}|^3} \left[ \frac{|\mu_{mi}|^2}{2|\omega_{mi}| Z} + 2i \frac{|\mu_{mi}|^2}{(2|\omega_{mi}| Z)^2} \right], \]
with the refractive index \( n(|\omega_{mi}|) = \sqrt{\epsilon(|\omega_{mi}|)} \). It is interesting to observe that, to leading order in \( Z \), only contributions due to the parallel component of the atomic dipole moment are contributing; contributions due to the perpendicular component of the atomic dipole moment appear only in next-to-leading order. Again, we need to take the real part of \( \Delta \mathcal{E}_{i}^{\text{ret}} \) to get the explicit form of the energy shift,
\[ \Delta \mathcal{E}_{i}^{\text{ret}} = \frac{1}{4\pi \epsilon_0} \sum_{m<i} \frac{|\omega_{mi}|^3}{|n(\omega_{mi}) + 1|^2} \times \left\{ \left[ |n(\omega_{mi})|^2 - 1 \right] \cos(2|\omega_{mi}| Z) - 2 \text{Im}[n(\omega_{mi})] \sin(2|\omega_{mi}| Z) + 2\left[ |n(\omega_{mi})|^2 - 1 \right] \cos(2|\omega_{mi}| Z) - 2 \text{Im}[n(\omega_{mi})] \sin(2|\omega_{mi}| Z) \right\}. \]

We see that, in the retarded regime, the two contributions to the shift of an excited state behave quite differently. The nonresidue contribution in Eq. (108) behaves as \( Z^{-3} \) (see the analysis of the ground-state shift in Sec. VII A), and the residue contribution in Eq. (130) depends on distance as \( Z^{-1} \). Although it would be tempting to jump to the conclusion that Eq. (130) always dominates, this might, in fact, not always be the case as the relative size of the two contributions also depends on the values of the dipole matrix elements involved, which can vary significantly. Furthermore, Eq. (130) is oscillatory so that, at least in principle, there are sets of parameters for which it vanishes. Finally, we remark that it is easy to verify that, in the limit of nonabsorptive dielectric media, our results reduce to those derived in Ref. [19].

VII. SPONTANEOUS DECAY RATES NEAR A HALF-SPACE

The spontaneous decay rates are given by the imaginary part of the complex self-energy, Eq. (110). As the nonresidue contributions to the self-energy in Eq. (108) are real, these contribute towards the energy-level shifts only, and the decay rates are contained solely in the residue contributions to the self-energy, Eq. (120), which are complex. In the nonretarded limit, the decay rates are given by the imaginary part of Eq. (126),
\[ \Delta \Gamma_{i}^{\text{nonret}} = \frac{1}{8\pi \epsilon_0 Z^3} \sum_{m<i} \left| \text{Im}[\epsilon(|\omega_{mi}|)] \right| |\epsilon(|\omega_{mi}|)| + 1 \left[ |\mu_{mi}|^2 + 2|\mu_{mi}^i|^2 \right], \]
and, in the retarded limit, by the imaginary part of Eq. (129),
\[ \Delta \Gamma_{i}^{\text{ret}} = -\frac{1}{2\pi \epsilon_0} \sum_{m<i} \frac{|\omega_{mi}|^3}{|n(\omega_{mi}) + 1|^2} \times \left\{ \left[ |n(\omega_{mi})|^2 - 1 \right] \sin(2|\omega_{mi}| Z) - 2 \text{Im}[n(\omega_{mi})] \cos(2|\omega_{mi}| Z) + 2\left[ |n(\omega_{mi})|^2 - 1 \right] \cos(2|\omega_{mi}| Z) - 2 \text{Im}[n(\omega_{mi})] \sin(2|\omega_{mi}| Z) \right\}. \]

The result in Eq. (131) is found to be in agreement with that derived in Ref. [18], their Eq. (128). A consistency check of Eq. (132) is that it reduces to the results given in Ref. [19] if we assume \( n(\omega) \) to be real and frequency independent.

As a numerical example, we plot the normalized lifetime of the atomic state \( |i \rangle \) decaying into a lower state \( |m \rangle \). For simplicity, we assume a two-level system and \( |\mu_{mi}| = 0 \) so that the atom is polarized horizontally with respect to the surface. Then, the normalized lifetime that we plot in Fig. 4 is given by
\[ \tau^{-1}_{||} = \frac{\Delta \Gamma_{i}^{\text{ret}}}{\Delta \Gamma_{i}^0}, \]
where the quantity \( \Delta \mathcal{E}_{i}^{\text{ret}} \) comes from Eq. (120) and \( \Delta \Gamma_{i}^0 \) is the well-known decay rate in free space,
\[ \Delta \Gamma_{i}^0 = \frac{|\omega_{mi}|^3 |\mu_{mi}|^2}{3\pi \epsilon_0 \hbar}. \]

VIII. SUMMARY AND CONCLUSIONS

We have shown that, starting from a gauge-independent microscopic model as represented by the Hamiltonian (19), it is possible to develop a formalism which allows calculating QED corrections in the presence of absorptive and dispersive boundaries. We have used a diagrammatic technique to integrate out the damped polaritons in order to arrive at a Dyson equation for the electromagnetic displacement-field propagator. We have solved this integral equation exactly using traceable methods. The knowledge of the exact propagator has enabled us to calculate analytically the one-loop self-energy diagram for an electron bound in an atom near a dielectric half-space and, hence, to determine its energy-level shifts and the change in transition rates, which derive from the real and imaginary parts of the electron’s self-energy, respectively. This serves as a proof of principle that the theoretical framework developed here works correctly and efficiently, as most of these results have, in one form or another, been derived previously by other methods, although often with considerably more effort or much less rigor, especially as regards basic principles.

We have looked at the role of the material’s absorption in some detail and have confirmed the previously known result that absorption has the most profound impact on the atomic system in the nonretarded regime, i.e., when the distance between the atom and the dielectric mirror is much smaller than the wavelength of the dominant atomic dipole transition. If the distance between the atom and the surface
far exceeds the wavelength of this dominant transition, then, to leading order, dispersion and absorption do not affect the ground-state shift for which only static polarizabilities matter. The next-to-leading order corrections are proportional to the damping constant of the Lorentz-type dielectric function, and it turns out that only the material’s absorption lines that lie in the low-frequency region have a significant impact on the ground-state energy-level shift. We have also re-derived the distance dependence of the excited energy-level shifts and spontaneous decay rates. We have confirmed the fact that, in the nonretarded regime, the absorption is of fundamental importance to both the change in decay rates and the energy-level shifts. For example, for an atom near a nondispersive dielectric, the spontaneous decay rate in the near zone comes out as a distance-independent constant \([19]\), whereas, in reality, when the absorption is taken into account, a distance dependence \(Z^{-1}\) is obtained. In the far zone or retarded limit, the presence of absorption does not affect the characteristic \(Z^{-1}\) behavior of the excited energy-level shift and spontaneous decay rates, even though the coefficients differ from the nondispersive case.

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**APPENDIX A: DRESSED PHOTON PROPAGATOR AS A BOUNDARY-VALUE PROBLEM**

We aim to show that the integral equation satisfied by the dressed photon propagator derived in Sec. IV B,

\[
D_{ik}(\mathbf{r}, \mathbf{r}'; \omega) = D_{ik}^{(0)}(\mathbf{r} - \mathbf{r}'; \omega) + \frac{K(\omega)}{\epsilon_0} \int d^3 r_1 g(r_1) D_{ij}^{(0)}(\mathbf{r} - \mathbf{r}_1; \omega) D_{jk}(\mathbf{r}_1, \mathbf{r}'; \omega)
\]

(A1)

can also be solved by considering it as a boundary-value problem. Recall that \(D_{ik}^{(0)}(\mathbf{r} - \mathbf{r}')\) is the photon propagator in free-space Eq. (60), and \(g(\mathbf{r})\) is a dimensionless coupling constant that is equal to unity in the region occupied by the dielectric and vanishes otherwise. To describe a dielectric half-space occupying the \(z < 0\) region of space as illustrated in Fig. 1, we take \(g(\mathbf{r}) = \theta(-z_1)\), where \(\theta\) is the Heaviside step function. Knowing that the free-space propagator satisfies the differential Eq. (39), we apply the same differential operator to Eq. (A1), and after a short calculation, we obtain the differential equation satisfied by the photon propagator in the half-space geometry,

\[
(\nabla_i \nabla_j - \delta_{ij} \nabla^2) \left[ 1 + \theta(-z) \frac{K(\omega)}{\epsilon_0} \right] D_{ij}(\mathbf{r}, \mathbf{r}'; \omega) - \omega^2 D_{ik}(\mathbf{r}, \mathbf{r}'; \omega) = \frac{\epsilon_0}{(2\pi)^3} \int d^3 q (q_i q_k - \delta_{ik} q^2) e^{i \mathbf{q} (\mathbf{r} - \mathbf{r}')}.
\]

(A2)
The RHS can be rewritten as
\[-\frac{1}{(2\pi)^3} \int d^3q (q^2 \delta_{ik} - q_i q_k) e^{iq \cdot (r - r')} = \nabla^2 \delta_{ik}^\perp (r - r'),\]
where \(\delta_{ik}^\perp (r - r')\) is the transverse \(\delta\) function. Now, it is more apparent that the RHS of Eq. (A2) is a distribution which, unlike the transverse \(\delta\) function, is sharply localized around the point \(r = r'\) because the nonlocal part of the transverse \(\delta\) function is removed by the application of the Laplacian as is obvious from the relation,
\[-\nabla^2 \left( \frac{1}{4\pi |r - r'|} \right) = \delta^{(3)}(r - r'). \quad (A3)\]
The locality of \(\nabla^2 \delta_{ik}^\perp (r - r')\) is very helpful towards the solution of the differential Eq. (A2), which is essentially a scattering problem. Its RHS contains a distribution, representing a pointlike source, and our task is to work out reflection and transmission at the boundary of the dielectric. In order to proceed any further, we need to specify physical situation, i.e., decide on which side of the boundary the source is placed. Since our ultimate aim is to work out the energy shift in an atom located outside the dielectric, we choose to consider the case \(z' > 0\). Then, we write Eq. (A2) in a piecewise manner; on the vacuum side, we have
\[(\nabla_i \nabla_j - \delta_{ij} \nabla^2) D_{ik}(r, r'; \omega) = \omega^2 D_{ik}(r, r'; \omega), \quad z > 0, \quad \text{(A4)}\]
and on the dielectric side, we have
\[(\nabla_i \nabla_j - \delta_{ij} \nabla^2) D_{ik}(r, r'; \omega) = \omega^2 D_{ik}(r, r'; \omega), \quad z < 0, \quad \text{(A5)}\]
with the behavior of the propagator \(D_{ik}(r, r'; \omega)\) across the interface, \(z = 0\) is still to be determined. The local character of the RHS of Eq. (A2) simplifies its solution in that it makes the RHS of Eq. (A5) go to zero. In order to solve Eqs. (A4) and (A5), we start with the following ansatz:
\[D_{ik}(r, r'; \omega) = \begin{cases} D_{ik}^{(0)}(r, r'; \omega), & z < 0, \\ D_{ik}^{(0)}(r, r' - r^0) + D_{ik}^{(r)}(r, r', \omega), & z > 0. \end{cases} \quad \text{(A6)}\]
On the vacuum side, we write the solution as a sum that consists of a particular solution \(D_{ik}^{(0)}(r, r'; \omega)\), which we already know from Sec. III B, Eq. (A2), and a solution \(D_{ik}^{(r)}(r, r'; \omega)\) of the corresponding homogeneous equation [i.e., Eq. (A4) with the RHS set to zero], which represents the correction due to reflection at the boundary. The solution on the dielectric side \(D_{ik}^{(0)}(r, r', \omega)\) represents the transmitted part and satisfies the homogeneous Eq. (A5). The homogeneous solutions \(D_{ik}^{(0)}(r, r', \omega)\) and \(D_{ik}^{(r)}(r, r', \omega)\) are chosen in such a way that the general solution in Eq. (A6) satisfies appropriate electromagnetic boundary conditions across the interface \(z = 0\). To see what these boundary conditions should be, recall the formal definition of the dressed propagator,
\[D_{ij}(r, r'; \omega) = -\frac{i}{\hbar} \text{Tr}[\hat{D}_i(r, t)\hat{D}_j(r', t')]|\Omega\rangle. \quad \text{(A7)}\]
The displacement operator \(D_i(r, t)\) satisfies Maxwell’s equations, which follow from the Heisenberg equations of motion for the field operators. Therefore, the photon propagator, by virtue of its definition (A7), when taken as a function of argument \(r\) and index \(i\), is required to satisfy Maxwell’s boundary conditions across the interface,
\[\begin{align*}
E_t \text{ continuous} & \quad \rightarrow \quad e^{-1}\hat{D}_{ij}|_{z=0^-} = \hat{D}_{ij}|_{z=0^+}, \\
\hat{D}_z \text{ continuous} & \quad \rightarrow \quad \hat{D}_{ij}|_{z=0^-} = \hat{D}_{ij}|_{z=0^+}, \quad \text{(A8)}
\end{align*}\]
with \(|\Omega\rangle = \{x, y\} \}

The apparent complication, arising from the appearance of a nonstandard distribution in the boundary-value problem (A2), is just an illusion. In fact, it is easier to find the solution of Eq. (A2) than it is to solve the differential equation satisfied by the Green’s function of the standard wave equation (see, e.g., Ref. [37]). Equations (A4) and (A5) together with the boundary conditions (A8) form a boundary-value problem, which is equivalent to the integral equation (A1) with the choices \(g(r) = \theta(-z_1)\) (dielectric occupying the left half-space) and \(z' > 0\) (source located in vacuum).

In the following, we will use Eq. (80) in the process of matching the boundary conditions. This is safe because we consider \(z = 0^\pm\), and the source located at \(z' > 0\) is always well away from the boundary so that \(z \neq z'\) is assured.

To proceed further, we note that taking the divergence of the integral equation in Eq. (A1) and using the fact that the free-space propagator is transverse \(\nabla_i D^{(0)}_{ik}(r - r'; \omega) = 0\), one infers that the dressed photon propagator is transverse everywhere as well,
\[\nabla_i D_{ik}(r, r'; \omega) = 0. \quad \text{(A9)}\]
With this, Eqs. (A4) and (A5) simplify further, and partially Fourier transformed into the \((q_1, z)\) space, cf. Eq. (65), they may be written as
\[\begin{align*}
\left(\nabla_z \cdot - q_1^2 + \omega^2\right) D_{ij}(z, z') & \quad = \omega^2 \delta_{ij}(q_1, z - z'), \quad z > 0, \quad \text{(A10)}
\left[\nabla_z \cdot - q_1^2 + \xi(\omega) q_2^2\right] D_{ij}(z, z') & \quad = 0, \quad z < 0, \quad \text{(A11)}
\end{align*}\]
where \(\delta_{ij}(q_1, z - z')\) is the Fourier transform of \(\delta_{ij}^\perp(r - r')\) with respect to \(r_\parallel - r'_\parallel\).
\[\delta_{ij}(q_1, z - z') = \frac{1}{2\pi^2} \int d^2(r_\parallel - r'_\parallel) e^{-i q_1 (r_\parallel - r'_\parallel)} \delta_{ij}^\perp(r - r'). \quad \text{(A12)}\]
The homogeneous solutions \(D_{ij}^{(r)}(r, r'; \omega)\) and \(D_{ij}^{(r)}(r, r', \omega)\) in Eq. (A6) must necessarily take the form
\[\begin{align*}
D_{ij}^{(r)}(z, z') & \quad = -\frac{i\epsilon_0}{2} \left[R_{ij}(z') e^{ik_z z} + S_{ij}(z') e^{-ik_z z}\right], \quad z > 0, \quad \text{(A12)}
D_{ij}^{(r)}(z, z') & \quad = -\frac{i\epsilon_0}{2} \left[T_{ij}(z') e^{-ik_z z} + U_{ij}(z') e^{ik_z z}\right], \quad z < 0, \quad \text{(A13)}
\end{align*}\]
with \(k_z = \sqrt{\omega^2 - q_1^2 + i\eta}\) and \(k_{z, d} = \sqrt{\xi(\omega) q_2^2 - q_1^2}\) and the square roots taken such that \(\text{Im}(k_z) \geq 0\) and \(\text{Im}(k_{z, d}) \geq 0\).
With this choice of sign for the square roots, the terms in Eqs. (A12) and (A13) that contain exponentials $e^{ikz}$ and $e^{ik_{zd}}$ are unphysical as they represent waves that diverge at infinity. Thus, we must set $S_j = 0 = U_{ij}$. The remaining two matrices $R_{ij}$ and $T_{ij}$ are determined by the requirement that Eq. (A6) satisfies the boundary conditions in Eq. (A8). We note that, in addition, the transversality of the dressed propagator, Eq. (A9), imposes rather stringent constraints on both $R_{ij}$ and $T_{ij}$. For example, the matrix $R_{ij}$ needs to be of the form

$$R_{ij} = v_i(q) r_j(q_1; z'),$$  \hspace{1cm} (A14)

where the vector $v$ is such that $q \cdot v = 0$, leading to

$$v = \left( v_x, v_y, -q_x v_y + q_y v_x / k_z \right),$$ \hspace{1cm} (A15)

with $q \equiv (q_1, k_z)$. One might pick $v_x = -q_y$ and $v_y = q_x$, so that

$$v = (-q_y, v_x, 0).$$ \hspace{1cm} (A16)

However, this choice is too restrictive on its own, as there is no a priori reason for $D_{ij}$ to vanish. Therefore, an additional basis vector is needed in order to span the amplitude $R_{ij}$ in full generality. An obvious and convenient choice is to choose a vector that is orthogonal to both $q$ and $v$,

$$w = v \times q = (q_x k_z - q_y k_z, -q_x^2).$$ \hspace{1cm} (A17)

Then, we can represent $R_{ij}$ as the linear combination,

$$R_{ij} = [av + \beta w]r_j(z'),$$

\hspace{1cm}\equiv e_i^T(k_z) r_j^{TE}(z') + e_i^T(k_z) r_j^{TM}(z'),

where we have recognized, apart from normalization factors, the transverse electric and transverse magnetic polarization vectors of Eq. (78). Similarly, we have

$$T_{ij} = e_i^T(-k_{zd}) h_j^{TE}(z') + e_i^T(-k_{zd}) h_j^{TM}(z').$$ \hspace{1cm} (A18)

We have chosen to write out the $k_z$ dependence of the polarization vectors, even though $k_z$ and $k_{zd}$ are expressible in terms of the frequency $\omega$ and the parallel wave vector $q_{\parallel}$ because this explicitly indicates the wave vector to which a given polarization vector is orthogonal. The decomposition into transverse electric and transverse magnetic components significantly simplifies the matching of boundary conditions. The dressed photon propagator can now be written in the form

$$D_{ij}(z,z') = e^{iq}(-k_{zd}q) e^{ik_{zd}} \Theta(-z)$$

$$+ \left[ e^i(k_z) r_j^{TE} e^{ikz} + \Theta(k_z e^i(k_z) e^{ik_{zd}} e^{-ik_{zd}(z-z')}) \right] \Theta(z).$$ \hspace{1cm} (A19)

The last term of which is the free-space photon propagator from Eq. (80) for $z - z' < 0$ as appropriate for the matching of boundary conditions at $z = 0$ when $z' > 0$. Imposing the boundary conditions of Eq. (A8), we find that

$$r_j^{z} = r_j e^i(-k_z) e^{i\omega^2/2k_z} e^{ikz},$$

$$t_j^{z} = t_j e^i(-k_z) e^{i\omega^2/2k_z} e^{ikz},$$

with $r_j$ and $t_j$ being the standard Fresnel’s reflection and transmission coefficients listed in Eq. (88). Plugging the above amplitudes into Eq. (A19), we then readily obtain the photon propagator given in Eq. (96).

We can readily apply the same methods to obtain the propagator in the case when the source is placed in the dielectric, i.e., for $z' < 0$. The calculation goes along exactly the same lines as for $z' > 0$, and one can show that the photon propagator, in the case of the source being placed in the dielectric, is given by

$$D_{ij}(r, r'; \omega) = \langle \Theta(-z)D^{(c)}_{ij}(r \rightarrow r'; \omega) - i\epsilon_0 \sum_k \int dq_1 \xi(\omega)\omega^2 2k_{zd} \rangle$$

$$\times e^{iq_{\parallel}(r - r') \cdot \xi(\omega) - i\omega(\xi(\omega)q_{\parallel} - k_{zd}q_1) r_{L}^z}$$

$$\times e^{-ik_{zd}(z - z')} + \Theta(z') \xi(\omega)q_{\parallel} - k_{zd}q_1 r_{L}^z e^{ik_{zd}(z - z')} \right].$$ \hspace{1cm} (A20)

Here, the reflection and transmission coefficients are those appropriate for left-incident modes; they are given by

$$r_{L}^{TE} = k_{zd} - k_z / k_z + k_{zd}, \quad r_{L}^{TM} = k_{zd} - \xi(\omega)k_z / \xi(\omega)k_z + k_{zd},$$ \hspace{1cm} (A21)

$$t_{L}^{TE} = 2k_{zd} / k_z + k_{zd}, \quad t_{L}^{TM} = 2k_{zd} / k_z + k_{zd}.$$ \hspace{1cm} (A21)

It is easily verified that $D_{ij}(r, r'; \omega)$ is, indeed, transverse everywhere.

**APPENDIX B: SIMPLE MODEL FOR $\epsilon(\omega)$**

In order to determine the dielectric permittivity of our model, we use the equations of motion for the fields that follow from the Hamiltonians (7)–(11) and the commutation relations (14)–(15),

$$\partial_t D(r,t) = 1 / \mu_0 \nabla \times B(r,t),$$ \hspace{1cm} (B1)

$$\partial_t B(r,t) = - \nabla \times E(r,t),$$ \hspace{1cm} (B2)

$$\partial_t X(r,t) = 1 / \lambda_0 \text{P}(r,t),$$ \hspace{1cm} (B3)

$$\partial_t P(r,t) = - M \omega_0^2 \text{X}(r,t) + g(r)E(r,t)$$

$$+ \int_0^\infty \rho_0 v^2 Y_v(r,t),$$ \hspace{1cm} (B4)

$$\partial_t Y_v(r,t) = 1 / \rho_0 Z_v(r,t),$$ \hspace{1cm} (B5)

$$\partial_t Z_v(r,t) = - \rho_0 v^2 Y_v(r,t) + \rho_0 v^2 X(r,t).$$ \hspace{1cm} (B6)

First, we deal with the subsystem consisting of the polarization field and the reservoir. It is well known [24] that, when a
quantized harmonic oscillator is coupled to a bath, its equation of
time takes the form of a quantum Langevin equation.
Thus, we expect the equation of motion for the polarization
field, which is nothing but a set of independent oscillators, to
take the form
\[ \mathcal{M} \frac{d^2}{dt^2} X(t) + \int_{-\infty}^{t} dt' \mu(t-t') \frac{d}{dt} X(t') + M \omega^2 X(t) = F_{\text{ran}}(t) + F_{\text{ext}}(t), \]
(B7)
where \( \mu(t-t') \) is the so-called memory function related
to dissipation and \( F_{\text{ran}}(t) \) represents some random force
operator (see, e.g., Ref. [24] for details). Both \( \mu(t-t') \) and
\( F_{\text{ran}}(t) \) arise as a consequence of the coupling to the bath
and are to be determined in terms of the parameters of our
model. The term \( F_{\text{ext}}(t) \) represents any external forces (i.e.,
those in addition to the harmonic restoring force) that may be
applied to the polarization field. To show that Eqs. (B3)–(B6)
indeed, combine to yield an equation of the form of Eq. (B7),
we eliminate \( P(t) \) and \( Z_v(t) \) and rewrite the equations for
\( X(t) \) and \( Y_v(t) \) as
\[
\left( \frac{\partial^2}{\partial t^2} + \omega^2 + \frac{1}{M} \int_0^{\infty} dv \rho_v v^2 \right) X(r, t) = \frac{1}{M} \int_0^{\infty} dv \rho_v v^2 Y_v(r, t),
\]
(B8)
\[
\left( \frac{\partial^2}{\partial t^2} + v^2 \right) Y_v(r, t) = v^2 X(r, t).
\]
(B9)
The most general solution of Eq. (B9) may be written as
\[ Y_v(r, t) = Y_v^H(r, t) + \int_{-\infty}^{t} dt' G_v(t-t') X(t'), \]
(B10)
where \( Y_v^H(r, t) \) is the solution of the homogeneous equation,
i.e., Eq. (B9) with its RHS set to zero,
\[ Y_v^H(r, t) = Y_v(r, 0) \cos(\nu t) + \frac{Z_v(r, 0)}{\rho_v} \sin(\nu t). \]
(B11)
We assume that bath operators \( Y_v(r, 0) \) and \( Z_v(r, 0) \) satisfy
the canonical commutation relations at the initial time \( t = 0 \),
cf. Eq. (15), which we take as a moment in the distant past
when the interaction has been switched on. The second term
in Eq. (B10) is a particular solution expressed in terms of the
Greens function,
\[ G_v(t-t') = \frac{1}{2\pi} \int_{-\infty}^{\infty} \frac{v^2}{\nu^2 - \omega^2 - i\epsilon} e^{-i\omega(t-t')} \]
(B12)
The \( i \epsilon \) prescription for handling the pole ensures that we have
a retarded Greens function with \( G_v(t-t') = 0 \) for \( t-t' < 0 \).
Equation (B12) is easily obtained from Eq. (B9) by using
Fourier transforms according to
\[ Y_v(r, \omega) = \int_{-\infty}^{\infty} dt e^{i\omega t} Y_v(r, t). \]
(B13)
Note that the choice of the retarded solution breaks
transformation invariance as has been noted in Ref. [24]. The integral
in Eq. (B12) is easily worked out using the residue theorem,
and Eq. (B10) may be rewritten as
\[ Y_v(r, t) = Y_v^H(r, t) + X(r, t) \]
\[ - \int_{-\infty}^{t} dt' \cos[\nu(t-t')] \frac{d}{dt} X(r, t), \]
(B14)
where we have integrated by parts. Plugging the above
equation into Eq. (B8), we obtain
\[ \mathcal{M} \frac{d^2}{dt^2} X(r, t) + \int_{-\infty}^{t} dt' \mu(t-t') \frac{d}{dt} X(r, t') + M \omega^2 X(r, t) \]
\[ = \int_0^{\infty} dv \rho_v v^2 Y_v^H(r, t), \]
(B15)
This is the quantum Langevin equation that follows from our
model. Comparing with Eq. (B7), lets us identify
\[ F_{\text{ran}}(t) = \int_0^{\infty} dv \rho_v v^2 Y_v^H(r, t), \]
(B16)
\[ \mu(t-t') = \int_0^{\infty} dv \rho_v v^2 \cos[\nu(t-t')]. \]
(B17)
Now, we are in the position to choose the bath oscillator masses \( \rho_v \). Having in mind a simple single-resonance model of the
dielectric permittivity, we choose \( \rho_v \) in such a way that the
friction term in Eq. (B15) is local in time, i.e., it is nonvanishing
only for \( t = t' \). This is achieved by choosing
\[ \rho_v = \frac{4\mathcal{M} \gamma}{\pi v^2}, \]
(B18)
which gives a frequency-independent coupling between the
bath and the polarization oscillators, cf. Eq. (10). Then,
Eq. (B15) becomes
\[ \frac{\partial^2}{\partial t^2} X(r, t) + 2\gamma \frac{\partial}{\partial t} X(r, t) + \omega^2 X(r, t) \]
\[ = \int_0^{\infty} dv \rho_v v^2 Y_v^H(r, t) + \frac{g(r)}{\mathcal{M}} E(r, t). \]
(B19)
We have augmented this equation by the “external force”
term that arises when the polarization field is coupled to the
electromagnetic field, which, according to Eqs. (B1) and (B2),
satisfies the equation of motion,
\[ \nabla \times [\nabla \times E(r, t)] + \mu_0 \sigma_0 \frac{\partial^2}{\partial t^2} E(r, t) = -\mu_0 g(r) \frac{\partial^2}{\partial t^2} X(r, t), \]
(B20)
with \( D(r, t) = \epsilon_0 E(r, t) + g(r) X(r, t) \). Similar to the reservoir
field discussed before, the most general solution of Eq. (B19)
exists as a sum of the homogeneous solution [i.e., the solution
of Eq. (B19) with the RHS set to zero and the assumption that
the oscillators are underdamped] and the particular solution.
The homogeneous solution is of the same form as Eq. (B11)
except for an additional damping factor proportional to \( e^{-\gamma t} \).
Since we assume that the initial time is a moment in the
distant past, we may discard the homogeneous solution, which
is exponentially small for \( \gamma t \gg 1 \). The particular solution is
so easily obtained in Fourier space and is given by
\[ X(r, t) = \frac{1}{2\pi \mathcal{M}} \int_{-\infty}^{\infty} d\omega \frac{F_{\text{ran}}(r, \omega) + g(r) E(r, \omega)}{\omega^2 - \omega^2 + 2i\gamma \omega} e^{-i\omega t}, \]
(B21)
where \( F_{\text{ran}}(r, \omega) \) is the Fourier transform of Eq. (B16) and
is given explicitly by
\[ F_{\text{ran}}(r, |\omega|) = 4\gamma Y_{|\omega|}(r, 0) + i \frac{\pi}{\mathcal{M}} |\omega| Z_{\omega}(r, 0). \]
Substitution of the solution (B21) into Eq. (B20) yields
\[ V \times [V \times E(r, \omega)] = -\mu_0 \varepsilon_0 \omega^2 \left[ 1 + \frac{g^2(r)}{M \varepsilon_0 \omega^2 - \omega^2 - 2i\gamma \omega} \right] E(r, \omega) \]
\[ = \mu_0 \frac{g(r)}{M} \frac{F_{\text{ran}}(r, \omega)}{\omega^2 - \omega^2 - 2i\gamma \omega}. \]  
(B23)

We may now read off the dielectric function given by
\[ \varepsilon / \varepsilon_0 = 1 + \frac{g^2(r)}{\omega^2 - \omega^2 - 2i\gamma \omega}. \]  
(B24)

with \( \omega_0^2 \equiv 1/M\varepsilon_0 \). The quantity that appears on the RHS is proportional to the so-called noise-current operator, which is introduced \textit{ad hoc} in the phenomenological formulation of the quantum theory developed in Ref. [8]. In fact, we have
\[ \mathbf{J}_n(r, \omega) = -i \omega g(r) \frac{F_{\text{ran}}(r, \omega)}{M} \frac{1}{\omega^2 - \omega^2 - 2i\gamma \omega}. \]  
(B25)

Since the operator \( F_{\text{ran}}(r, \omega) \) depends only on the initial coordinates and momenta of the bath, cf. Eq. (B22), for which the commutation relations are known, it is relatively easy to verify that
\[ [J_i(r, \omega), J_j^+(r', \omega')] = 4\pi\hbar \varepsilon_0 \text{Im}[\varepsilon(r, \omega)] \omega^2 \delta^{(3)}(r - r') \delta(\omega - \omega') \delta_{ij}. \]  
(B26)

This derivation justifies these phenomenologically introduced commutation rules on a microscopic level. We just note that this result differs from Eq. (C2) to be used in the following Appendix by a factor of \((2\pi)^2\) due to a different definition of the Fourier transform.

### APPENDIX C: PHOTON PROPAGATOR FROM PHENOMENOLOGICAL QED

The phenomenological theory of quantum electrodynamics as developed in Ref. [8] gives the electric-field operator as
\[ E_i(r,t) = -i \mu_0 \int d^3 \mathbf{r} \int_0^\infty d\omega e^{-i\omega t} \times G_{ik}(r, \omega) J_j(r, \omega) + \text{H.c.}, \]  
(C1)

where \( J_j(r, \omega) \) is the so-called noise-current operator, satisfying the following commutation relation,
\[ [J_i(r, \omega), J_j^+(r', \omega')] = \frac{\hbar \varepsilon_0}{\pi} \text{Im}[\varepsilon(r, \omega)] \omega^2 \delta^{(3)}(r - r') \delta(\omega - \omega') \delta_{ij}, \]  
(C2)

and \( G_{ik}(\mathbf{r}, \omega) \) is the Green’s function of the wave equation satisfying
\[ (\nabla_i \nabla_j - \delta_{ij} \nabla^2) G_{jk}(\mathbf{r}, \mathbf{r}', \omega) = -\varepsilon(\mathbf{r}, \omega) \omega^2 G_{ik}(\mathbf{r}, \mathbf{r}', \omega) \]
\[ = \delta_{ik} \delta^{(3)}(\mathbf{r} - \mathbf{r}'), \]  
(C3)

with the additional requirement that it is retarded in time. Note, however, that there is no transversality condition imposed, and the RHS of Eq. (C3) is just a diagonal \( \delta \) function. For an overview of the noise-current approach and some applications, see Ref. [38]. In the following, we will use two properties of the Green’s tensor, in particular, its reciprocity,
\[ G_{ik}(\mathbf{r}, \mathbf{r}', \omega) = G_{ki}(\mathbf{r}', \mathbf{r}, \omega), \]  
(C4)

and the integral relation,
\[ \int d^3 \mathbf{r} \omega^2 \text{Im}[G(\mathbf{r}, \omega)] G^*_j(\mathbf{r}, \mathbf{r}', \omega) G_{jk}(\mathbf{r}, \mathbf{r}', \omega) = \text{Im}[G_{ij}(\mathbf{r}, \mathbf{r}', \omega)]. \]  
(C5)

To prove the latter, one multiplies Eq. (C3) from the left by \( G^*_j(\mathbf{r}, \mathbf{r}', \omega) \) and integrates over \( \mathbf{r} \). Then, taking the difference between the resulting relation and its complex conjugate integrated by parts yields Eq. (C5).

In order to calculate the Feynman propagator of the electric-field operator, i.e., the quantity,
\[ D_{ij}^E(\mathbf{r}, \mathbf{r}', t, t') = \frac{-i}{\hbar} \langle 0 | T[E_i(\mathbf{r}, t) E_j(\mathbf{r}', t')] | 0 \rangle, \]  
(C6)

we substitute, into the above definition, the operator (C1) and use Eqs. (C2)–(C5). We arrive at
\[ D_{ij}^E(\mathbf{r}, \mathbf{r}', t, t') = \frac{-i}{\pi \varepsilon_0} \int_0^\infty d\omega \omega^2 \theta(t - t') e^{-i\omega(t-t')} \]
\[ + (\theta(t' - t) e^{i\omega(t'-t)}) \text{Im}[G_{ij}(\mathbf{r}, \mathbf{r}', \omega)]. \]

Now, we carry out the Fourier transform with respect to \( t - t' \) using the distributional identities,
\[ \int_0^\infty d\tau e^{\pm i\Omega \tau} = \pi \delta(\Omega) \pm \frac{i}{\Omega}, \]  
(C7)

where \( \Omega \) denotes the Cauchy principal value and obtain
\[ D_{ij}^E(\mathbf{r}, \mathbf{r}', \Omega) = \frac{2}{\pi \varepsilon_0} \int_0^\infty d\omega \omega^3 \text{Im}[G_{ij}(\mathbf{r}, \omega)] \]
\[ - \frac{i}{\pi \varepsilon_0} \int_0^\infty d\omega \omega^2 \delta(\Omega - \omega) + \delta(\Omega + \omega) \]
\[ \times \text{Im}[G_{ij}(\mathbf{r}, \omega)]. \]  
(C8)

The Green’s tensor must satisfy retarded boundary conditions in time in order to preserve causality. This means that it is analytical in the upper half of the complex \( \omega \) plane. Analyticity in the upper half of the plane leads to Kramers-Kronig relations [39] so that the Green’s tensor inherits the causality properties of the permittivity. In particular, its imaginary part is an odd function of frequency \( \omega \) whereas, its real part is even in \( \omega \). With that, we can proceed to deal with the principal-value integral in Eq. (C8). Since \( \text{Im}[G_{ij}(\mathbf{r}, \mathbf{r}', -\omega)] = -\text{Im}[G_{ij}(\mathbf{r}, \mathbf{r}', \omega)] \) and the remaining part of the integrand is also odd, we extend the lower integration limit to \(-\infty\) and compensate by multiplying by 1/2. On the other hand, the real part of the Green’s tensor is even in \( \omega \) so that we can replace
\[ \text{Im}[G_{ij}(\mathbf{r}, \mathbf{r}', \omega)] \rightarrow \frac{1}{i} G_{ij}(\mathbf{r}, \mathbf{r}', \omega), \]  
(C9)

without changing the value of the integral. Thus, the principal-value integral in Eq. (C8) becomes
\[ \frac{\text{P}}{i\pi \varepsilon_0} \int_0^\infty d\omega \frac{\omega^3}{\Omega^2 - \omega^2} G_{ij}(\mathbf{r}, \mathbf{r}', \omega). \]  
(C10)

To work out this integral, we consider a contour of integration \( \gamma \) that runs from \(-\infty\) to \(\infty\) and above the poles at \(\omega = \pm \Omega\) and
then closes up in the upper half of the $\omega$ plane along the large semicircle $|\omega| \to \infty$. Because the Green’s tensor is analytic in the upper half plane, the such-calculated integral vanishes, and we can express the principal-value integral as

$$ P \int = - \int_{\gamma^-} - \int_{\gamma^+} - \int_{\Gamma}, $$

where $\gamma^\pm$ denotes the clockwise contours that go around the poles at $\omega = \pm \Omega$, respectively, and $\Gamma$ denotes the contribution from the large semicircle taken counterclockwise. Using the residue theorem, we derive that the contribution from $\gamma^+$ is given by

$$ - \frac{1}{\epsilon_0} \Omega^2 G_{ij}(r, r'; \Omega). $$

The large semicircle $\Gamma$ contributes the $\delta$ function,

$$ - \frac{1}{\epsilon_0} \delta_{ij} \delta(3)(r - r') $$

for whose calculation we have used the fact that, asymptotically, the Green’s tensor behaves as [38]

$$ \lim_{|\omega| \to \infty} \omega^3 G_{ij}(r, \omega) = - \delta_{ij} \delta(3)(r - r'). $$

The $\delta$-function integral in Eq. (C8) is easily seen to be

$$ - i \frac{\Omega^2}{\epsilon_0} \text{Im}[G_{ij}(r, r'; |\Omega|)], $$

so that the final result for the relation between the photon propagator and the Green’s function of the wave equation on the real $\Omega$ axis can be compactly written as

$$ D_{ij}^R(r, r'; \Omega) = - \frac{\Omega^2}{\epsilon_0} G_{ij}(r, r'; |\Omega|) - \frac{1}{\epsilon_0} \delta_{ij} \delta(3)(r - r'). $$

A similar formula has been given in Ref. [40]. We would like to use this result for a comparison with the results of Sec. IV B.

First, we need to emphasize that what we have calculated here is the propagator for the electric field $E$, whereas, Sec. IV B derives the propagator for the displacement field $D$. Therefore, the results can coincide only when $r$ and $r'$ are both located outside the dielectric, which is why we restrict ourselves to this case. Then, the Green’s tensor $G_{ij}(r, r'; \omega)$ splits into a free-space part $G_{ij}^{(0)}$ and a correction $G_{ij}^{(r)}$ that describes the reflection of the electromagnetic field from the surface, and Eq. (C16) can be rewritten as

$$ D_{ij}^{E}(r, r'; \Omega) = - \frac{\Omega^2}{\epsilon_0} G_{ij}^{(0)}(r - r'; |\Omega|) + \delta_{ij} \delta(3)(r - r'). $$

This makes clear that the Feynman propagator is an even function of $\Omega$, unlike the Green’s function of the wave equation, which has the same analytical structure as the dielectric function. It is not difficult to verify that, for the particular geometry considered here, the dielectric half-space Eq. (C17), indeed, holds. The terms in square brackets combine to deliver the transverse free-space propagator as given in Eq. (42). The reflected part $G_{ij}^{(r)}(r, r'; |\Omega|)$, which can be found, e.g., in Ref. [8], satisfies the homogeneous wave equation. Therefore, it is automatically transverse,

$$ \nabla_i G_{ij}^{(r)}(r - r'; \omega) = 0, $$

and for real $\omega$, it coincides with the reflected part of the photon propagator $D_{ij}^{E}(r, r'; \omega)$ given in Eq. (96), although away from the real axis, they are different due to the different boundary conditions in time. $D_{ij}^{E}(r, r'; t - t')$ gives the retarded solutions of the wave equation.

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[22] All the quantum fields here and in the following are, of course, time dependent, although we suppress this time dependence in the arguments for brevity.
[27] At this stage, the analogy between the quantity $(\epsilon_0 M)^{-1}$ and the plasma frequency as commonly introduced in the free-electron model of a metal is not readily apparent. This choice of notation justifies itself once the precise form of $\rho_\omega$ in Eqs. (17) and (18).
is chosen so that the dielectric function of the model emerges
(see Appendix B for details).